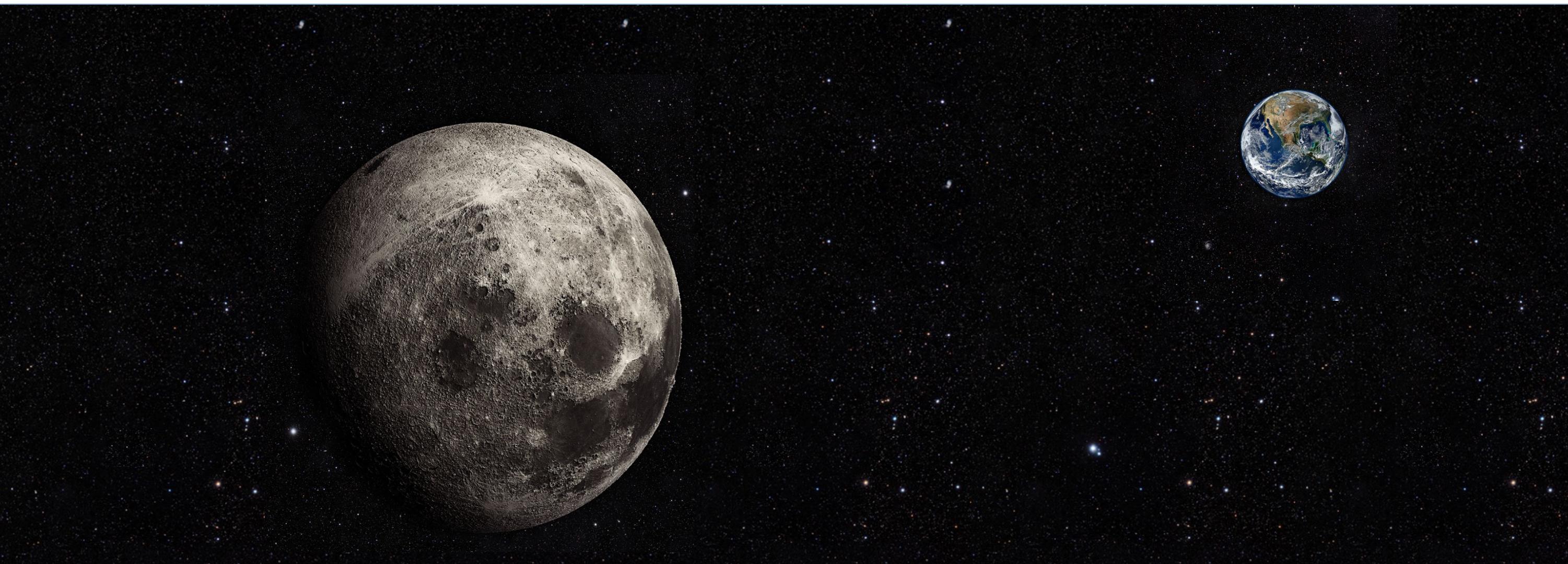


Comparative Analysis and Design of a Dual-Satellite System for Lunar Rover Localization



KAILA M. Y. COIMBRA AND GRACE GAO



A New Era in Lunar Exploration

*Dates are from launch to landing.

“Success” indicates successful landing.

Chandrayaan-3 (success)^[2]

Jul 14 – Aug 23, 2023

ISRO’s lander and rover (IND)

iSpace M1 (failed)

Dec 11, 2022 – Apr 25, 2023

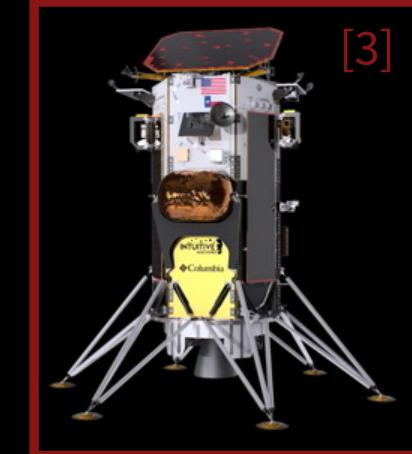
iSpace’s lunar lander;
crashed during landing (JPN)



Peregrine Mission 1 (failed)

Jan 8 – Jan 18, 2024

Astrobotic’s lunar
landing mission,
propellant leak (USA)



IM-1 (success)^[3]

Feb 15 – Feb 22, 2024

Intuitive Machines’s
Nova-C Odysseus
lander (USA)



Artemis I (success)

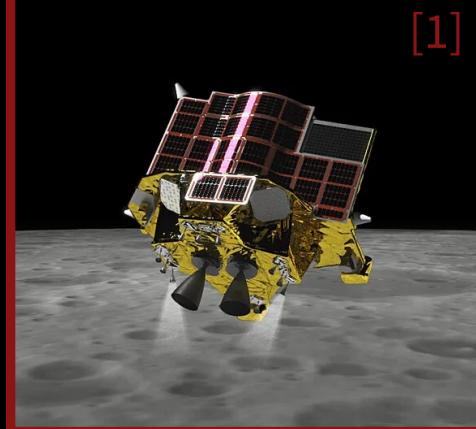
Nov 16 – Dec 11, 2022

NASA’s uncrewed test flight (USA)

Luna 25 (failed)

Aug 10 – Aug 19, 2023

Roscosmos’s lunar landing mission;
crashed into surface (RUS)



SLIM (success)

Sep 6, 2023 – Jan 19, 2024

JAXA’s precision lunar lander (JPN)

Chang’e-6 (success)^[4]

May 3 – Jun 1, 2024

Brought back first samples from
far side of the Moon (CHN)

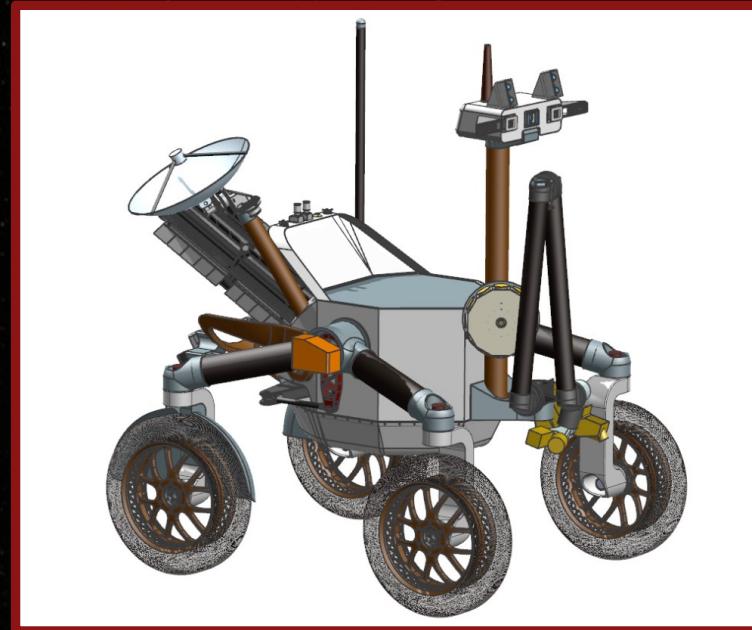
DRO-A and DRO-B (failed)

Mar 13 – Mar 13, 2024

Pair of lunar navigation satellites,
lost during launch (CHN)

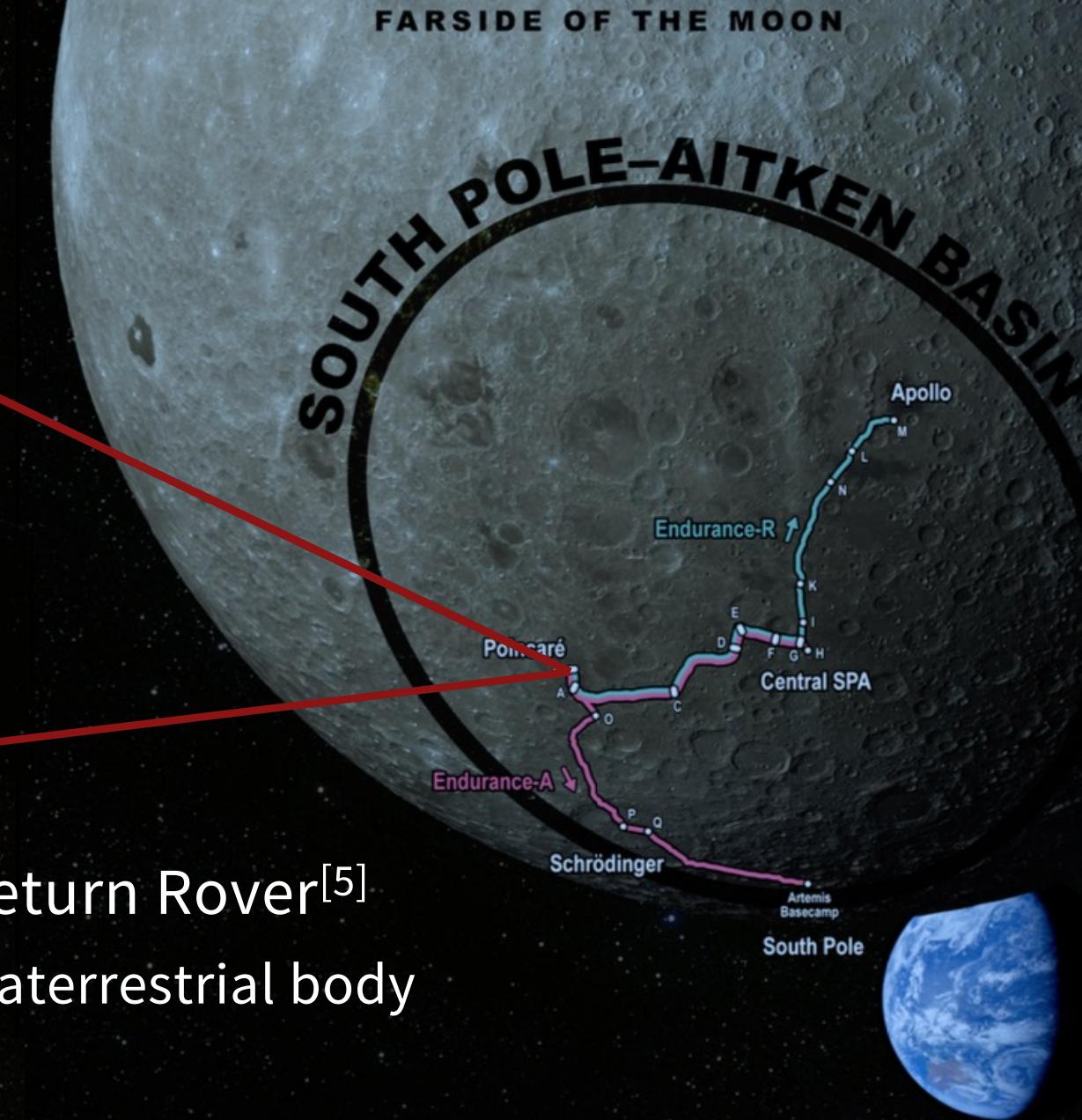
NASA Endurance Mission

FARSIDE OF THE MOON



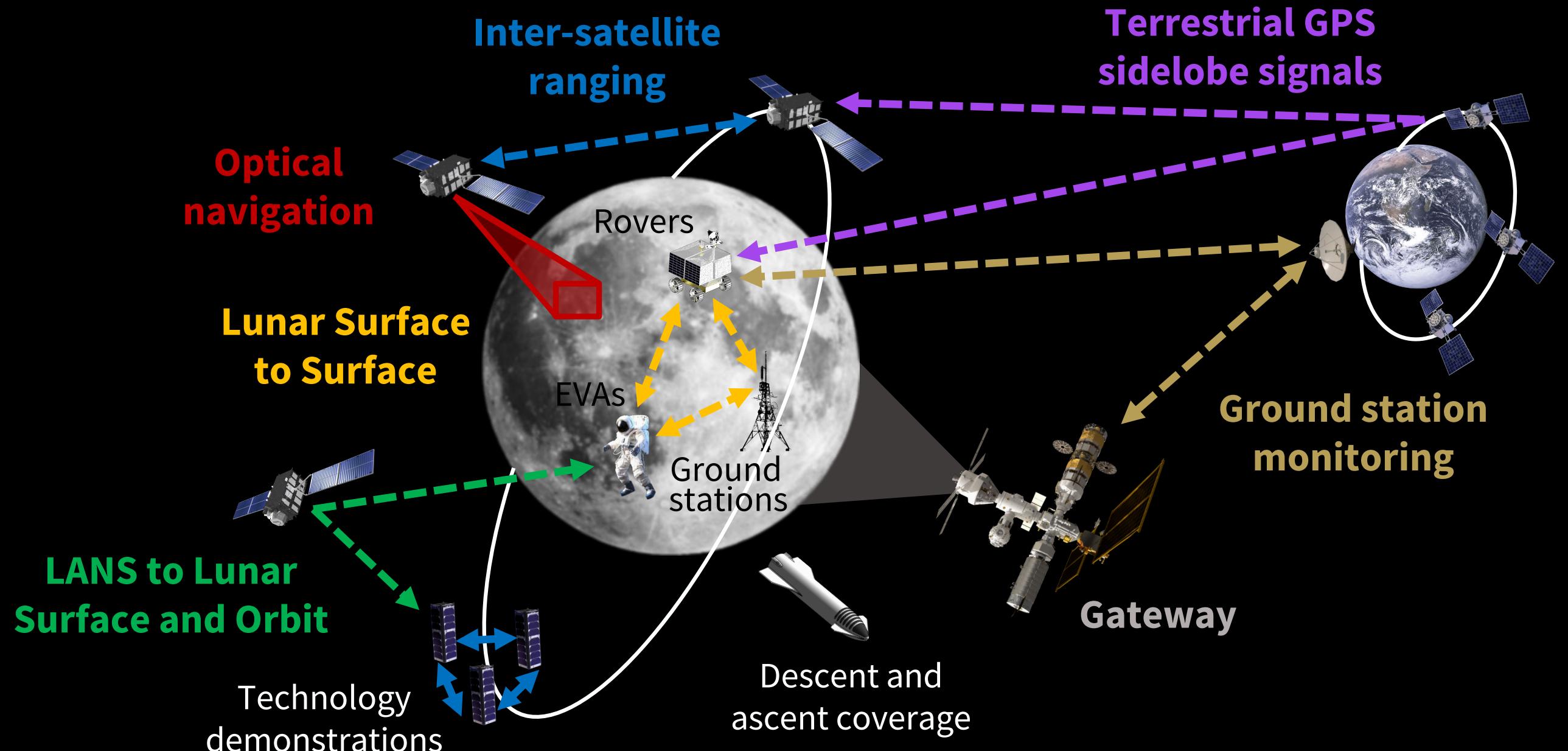
Lunar SPA Basin Traverse and Sample Return Rover^[5]

- First rover to traverse 2000 km on an extraterrestrial body
- *Launch:* ~2032
- *Mission:* Collect 12 samples along its traverse and return samples to Artemis astronauts



Earth image: Apollo 11 / NASA / JSC.
Moon topography: LRO LOLA.
Moon image mosaic: LRO WAC / LRO LOLA /
NASA's Scientific Visualization Studio.
Sky: Taurus / NSF NOIRLab / Akira Fujii.
Composed by James Tuttle Keane.

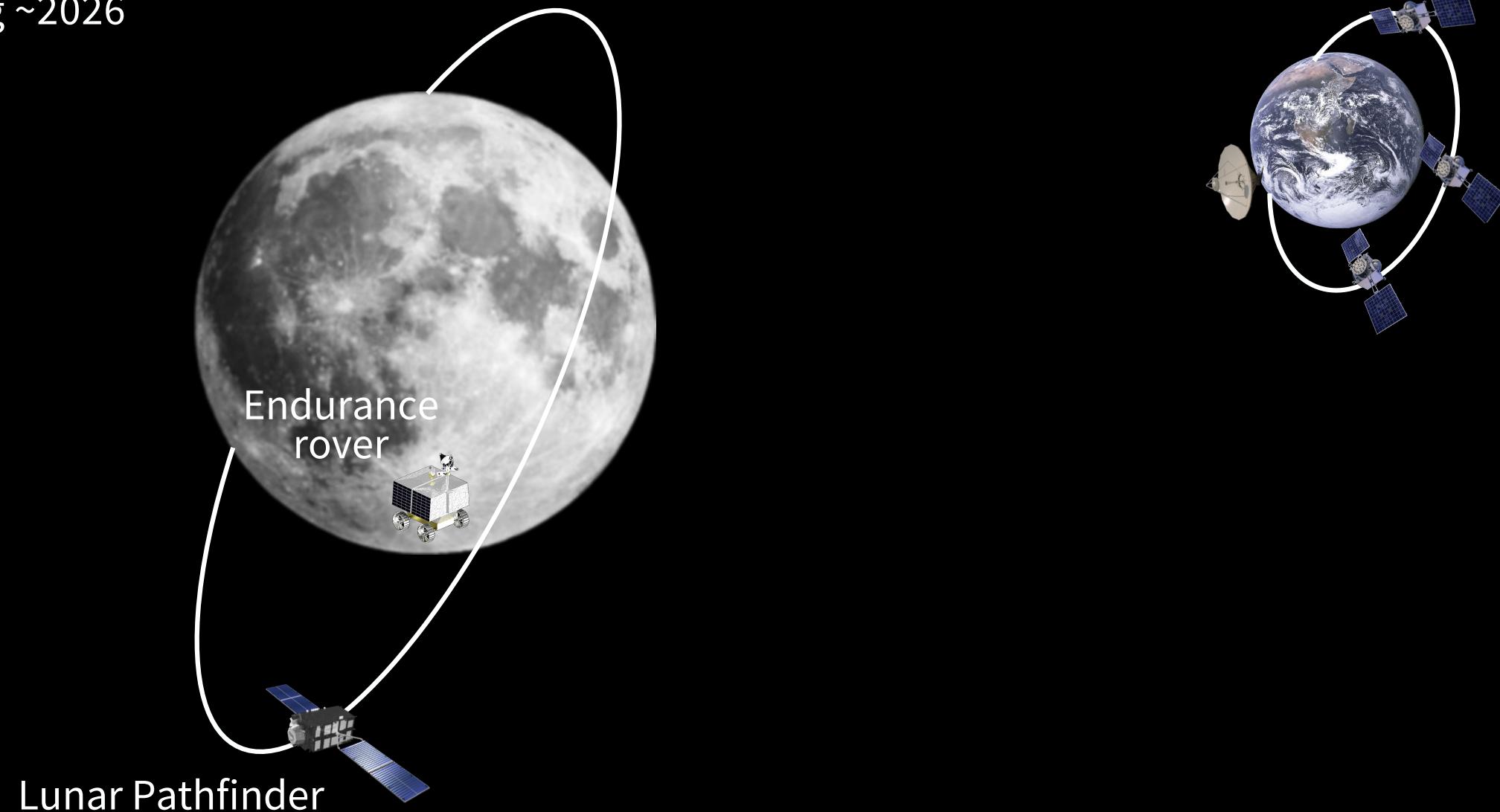
The Upcoming Lunar PNT Landscape



Lunar Pathfinder

Data Relay Satellite in an elliptical lunar frozen orbit (ELFO)^[6]

- *Mission:* Provide communication services to any lunar asset
- Operational starting ~2026

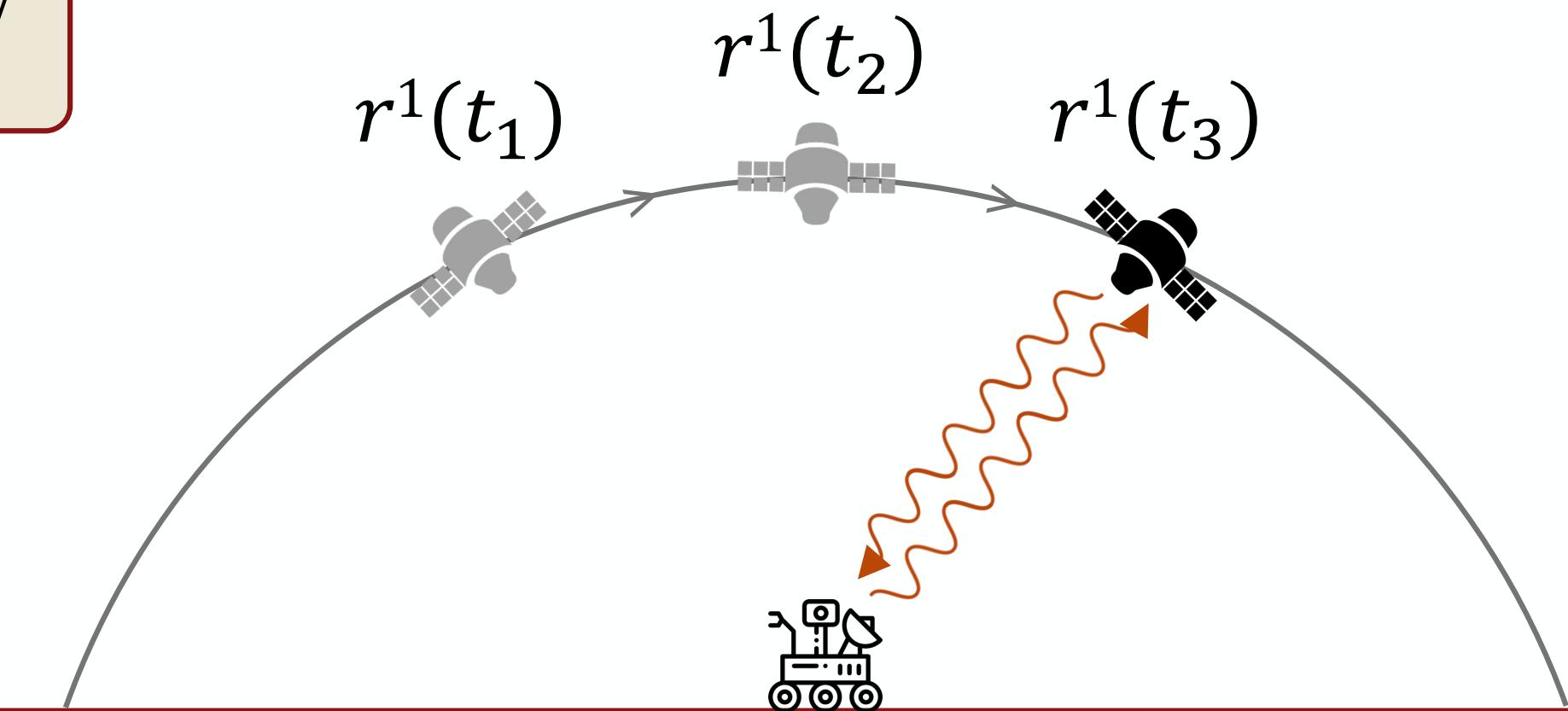


Prior Work: How to localize with only one satellite?

Approach: The rover, while stationary, accumulates navigation observables from a single satellite and refines its position estimate over time.

(1) Two-way ranging measurements^[7]

- Rover needs to be stationary for up to 4 hours to achieve sub-10 m accuracy
- Requires a navigation payload



Prior Work: How to localize with only one satellite?

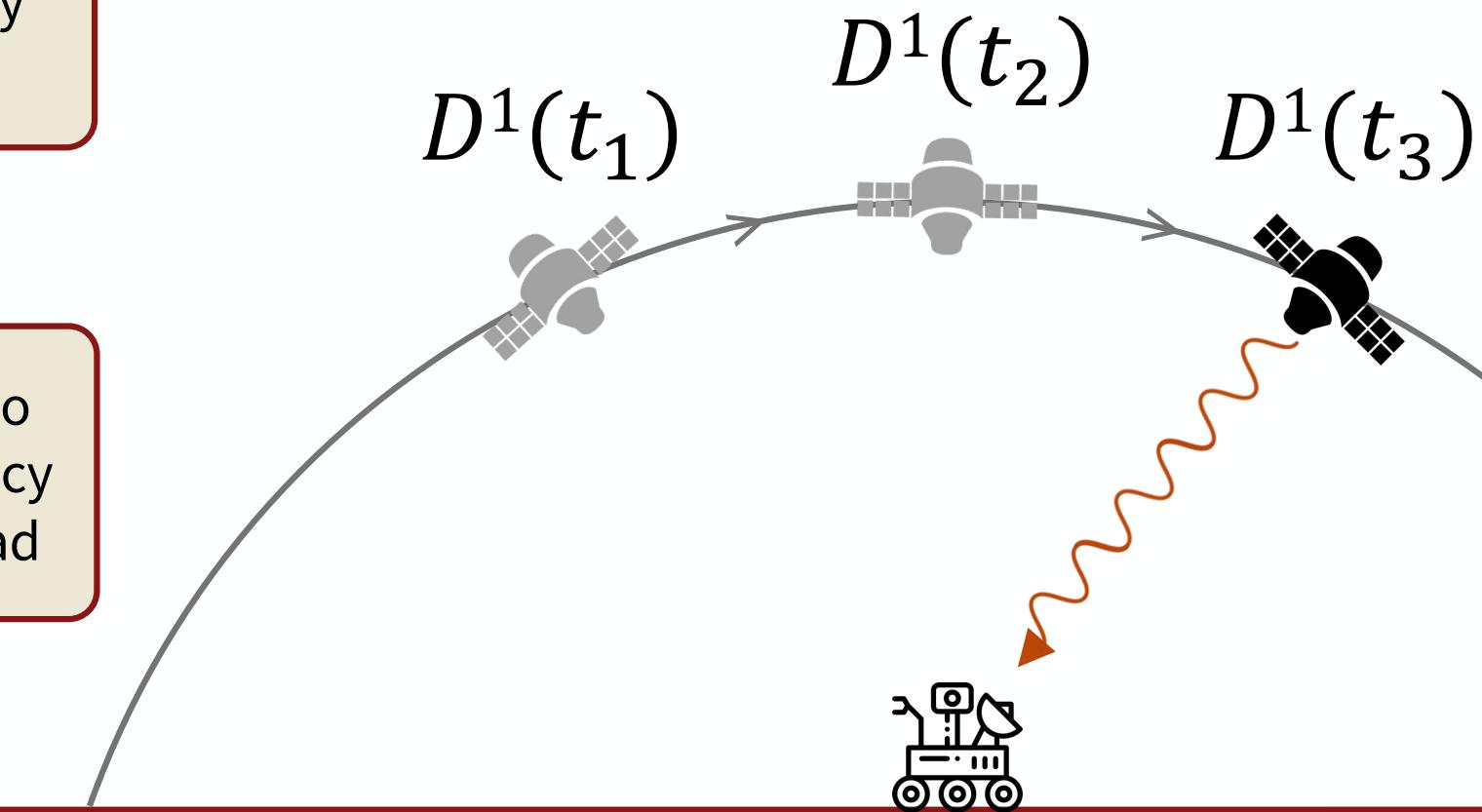
Approach: The rover, while stationary, accumulates navigation observables from a single satellite and refines its position estimate over time.

(1) Two-way ranging measurements^[7]

- Rover needs to be stationary for up to 4 hours to achieve sub-10 m accuracy
- Requires a navigation payload

(2) Doppler shift measurements^[8]

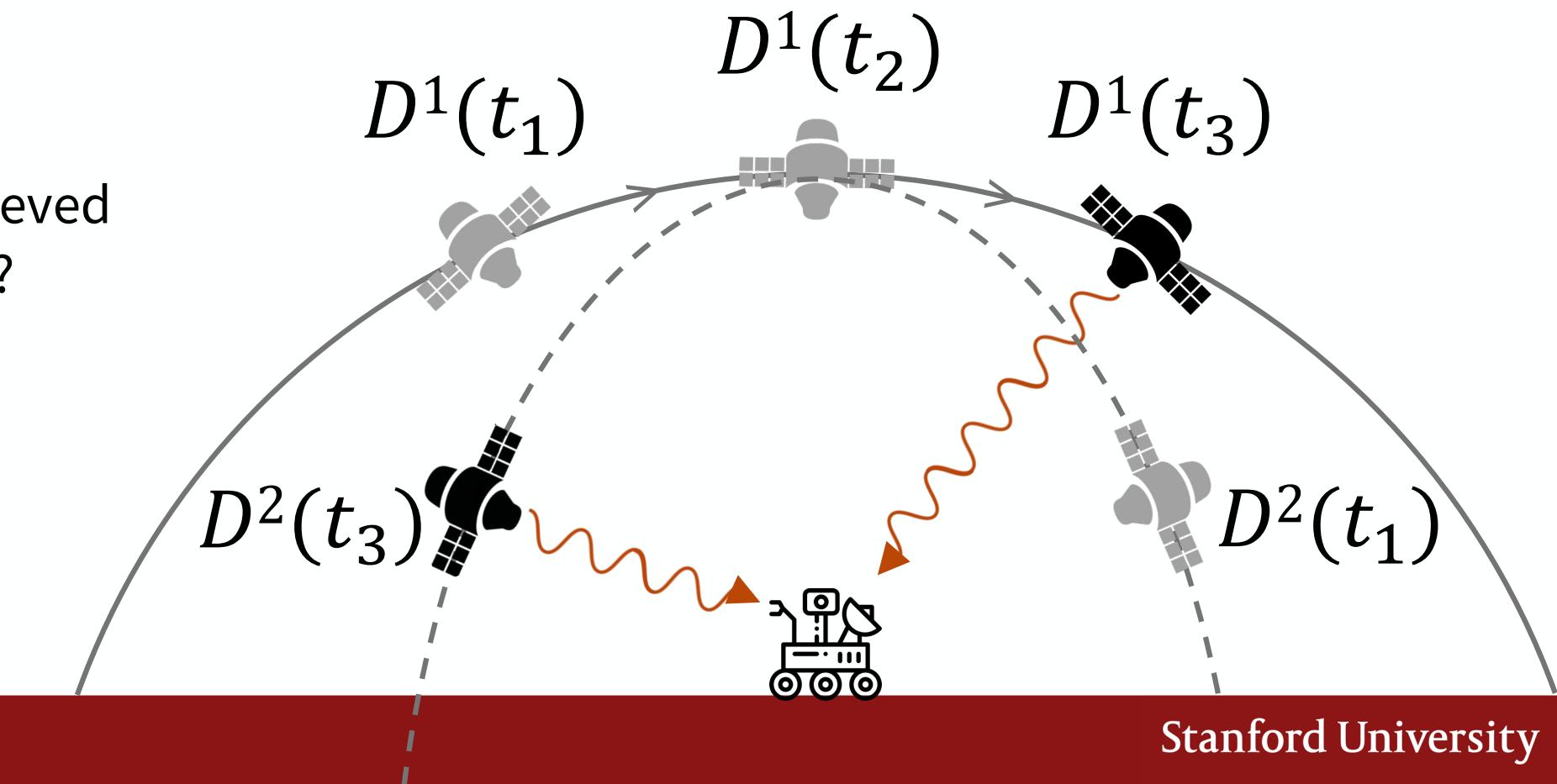
- Rover needs to be stationary for up to 16 hours to achieve sub-10 m accuracy
- Does not require a navigation payload

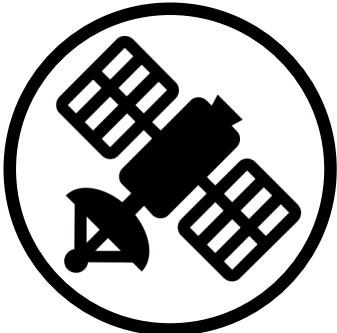


Problem Scenario: Introducing a second satellite

Key Questions:

- (1) What will the orbit of the auxiliary satellite have to look like to best complement the Lunar Pathfinder's coverage?
- (2) What *improvement* in the rover's localization accuracy can be achieved by introducing a second satellite?





Measurement Modeling and Filtering

- Simulate and filter realistic measurements from the satellites' communication signals to obtain a positioning fix.



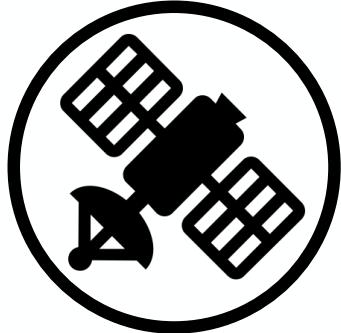
Orbit Design for Auxiliary Satellite

- Determine the orbital elements for the auxiliary satellite that maximizes coverage of the lunar South Pole region.



State Estimation Performance and Comparison

- Evaluate the rover's absolute localization accuracy using two non-navigation satellites and compare to the single-satellite scenario.



Measurement Modeling and Filtering

- Simulate and filter realistic measurements from the satellites' communication signals to obtain a positioning fix.



Orbit Design for Auxiliary Satellite

- Determine the orbital elements for the auxiliary satellite that maximizes coverage of the lunar South Pole region.



State Estimation Performance and Comparison

- Evaluate the rover's absolute localization accuracy using two non-navigation satellites and compare to the single-satellite scenario.

Measurement and Error Modeling

Determine pseudorange rate measurements from Doppler shift.

$$\dot{\rho} = -\frac{Dc}{f_C}$$



In simulation, generate pseudorange rate measurements that the rover will observe.

$$\tilde{\dot{\rho}}(t) = \mathbf{v}_s(t) \cdot \frac{\mathbf{r}_s(t) - \mathbf{r}_r}{\|\mathbf{r}_s(t) - \mathbf{r}_r\|} + c(\dot{\delta t}_r - \dot{\delta t}_s) + \epsilon_{\dot{\rho}}$$

True range rate Relative clock drift Error

Measurement and Error Modeling

Determine pseudorange rate measurements from Doppler shift.

$$\dot{\rho} = -\frac{Dc}{f_C}$$



In simulation, generate pseudorange rate measurements that the rover will observe.

$$\tilde{\dot{\rho}}(t) = \mathbf{v}_s(t) \cdot \frac{\mathbf{r}_s(t) - \mathbf{r}_r}{\|\mathbf{r}_s(t) - \mathbf{r}_r\|} + c(\dot{\delta t}_r - \dot{\delta t}_s) + \epsilon_{\dot{\rho}}$$

True range rate Relative clock drift Error

Measurement Errors

$$\epsilon_{\dot{\rho}} \sim \mathcal{N}(0, \sigma_{\dot{\rho}}^2)$$

Contributions include thermal and phase noise (phase scintillation is negligible).

$$\sigma_{\dot{\rho}}^2 = \sigma_t^2 + \sigma_c^2$$

Measurement and Error Modeling

Determine pseudorange rate measurements from Doppler shift.

$$\dot{\rho} = -\frac{Dc}{f_C}$$



In simulation, generate pseudorange rate measurements that the rover will observe.

$$\tilde{\dot{\rho}}(t) = \mathbf{v}_s(t) \cdot \frac{\mathbf{r}_s(t) - \mathbf{r}_r}{\|\mathbf{r}_s(t) - \mathbf{r}_r\|} + c(\dot{\delta t}_r - \dot{\delta t}_s) + \epsilon_{\dot{\rho}}$$

True range rate Relative clock drift Error

Measurement Errors

$$\epsilon_{\dot{\rho}} \sim \mathcal{N}(0, \sigma_{\dot{\rho}}^2)$$

Contributions include thermal and phase noise (phase scintillation is negligible).

$$\sigma_{\dot{\rho}}^2 = \sigma_t^2 + \sigma_c^2$$

Satellite Ephemeris Errors

The rover has erroneous knowledge of the satellite's position and velocity when predicting its expected measurements for filtering.

$$\tilde{\mathbf{r}}_s = \mathbf{r}_s + \epsilon_{e,p} \quad , \quad \epsilon_{e,p} \sim \mathcal{N}(0, \sigma_{e,p_{xyz}}^2 \mathbf{I}_{3 \times 3})$$

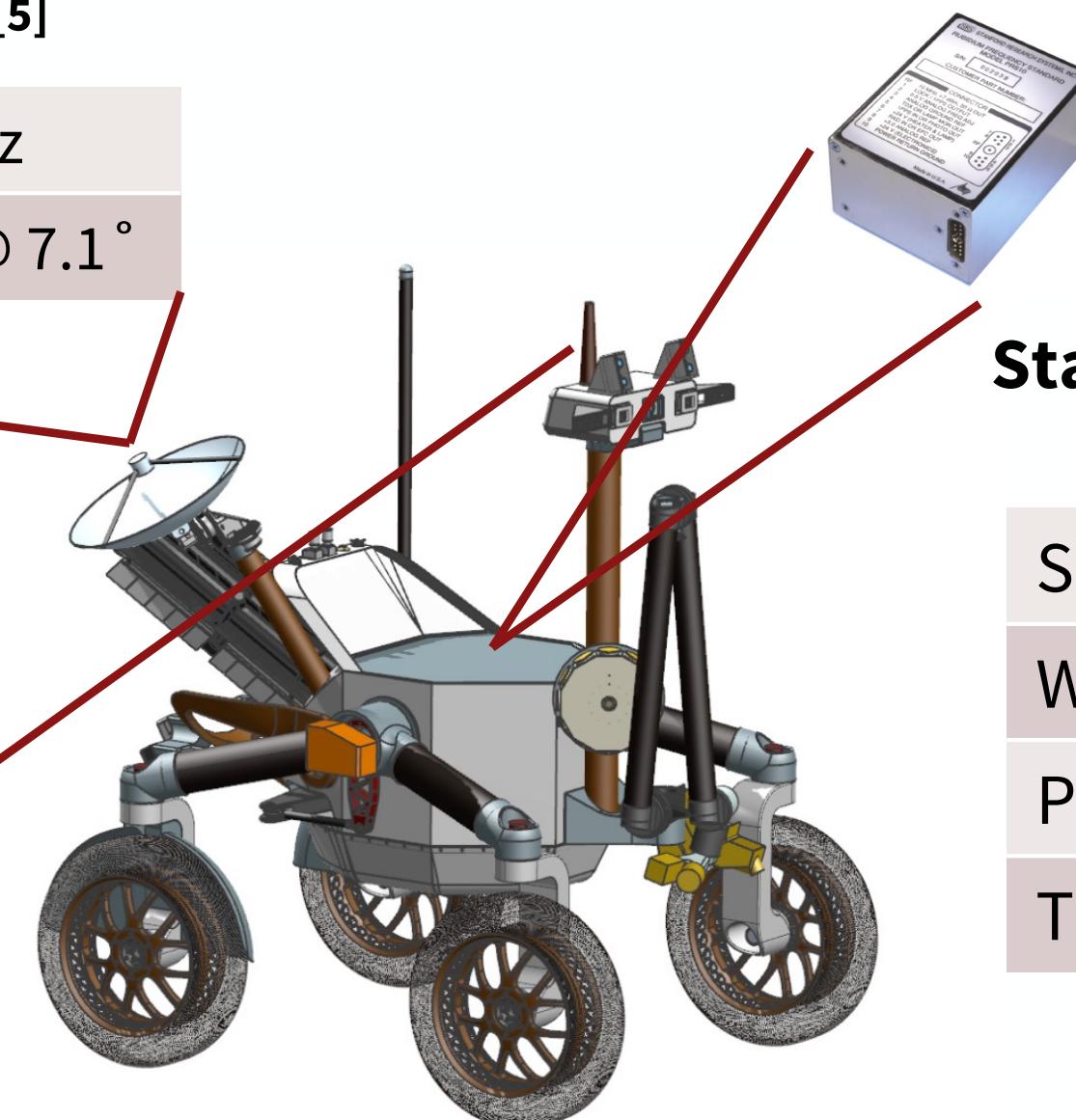
$$\tilde{\mathbf{v}}_s = \mathbf{v}_s + \epsilon_{e,v} \quad , \quad \epsilon_{e,v} \sim \mathcal{N}(0, \sigma_{e,v_{xyz}}^2 \mathbf{I}_{3 \times 3})$$

Endurance Rover's Relevant Components

75 cm high gain omnidirectional antenna^[5]

S-band frequency ^[4]	2050 Hz
EIRP @ HPBW ^[4]	26.5 dBW @ 7.1 °

Low gain antenna

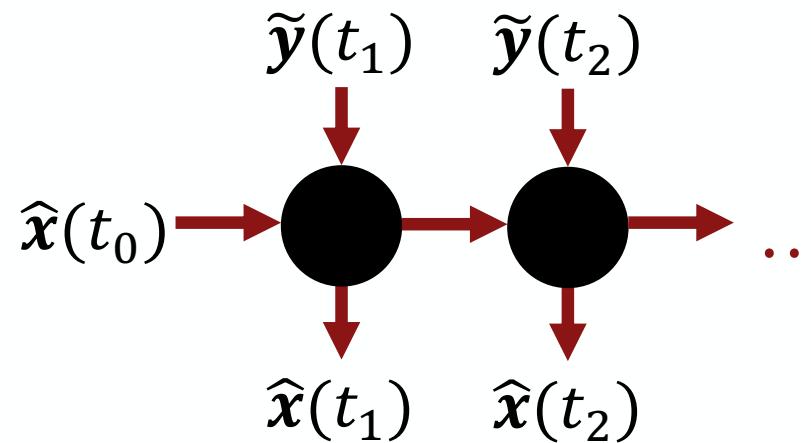


Stanford Research Systems PRS 10 clock^[9, 10]

Size	155 cm ³
Weight	0.6 kg
Power	14.4 W
TDEV per day	0.07 µs

State Estimation Considerations

Extended Kalman Filter



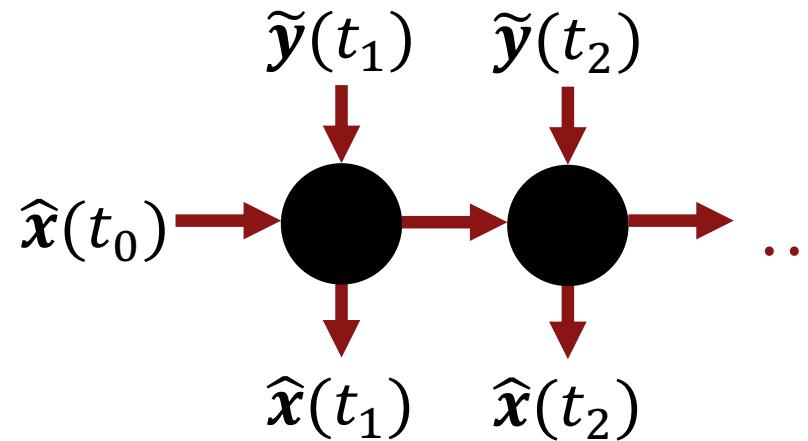
EKF processes one measurement at each time step.

- Not desirable for system with limited observability
- Consider filter that processes multiple measurements over time

	Rover state	Measurements
Observed	\tilde{x}	\tilde{y}
Expected	\hat{x}	\hat{y}

State Estimation Considerations

Extended Kalman Filter

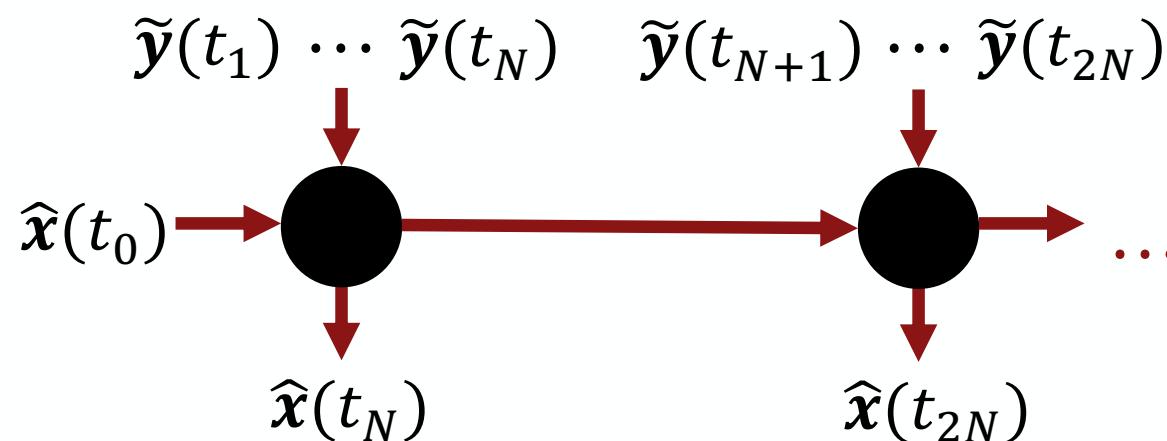


EKF processes one measurement at each time step.

- Not desirable for system with limited observability
- Consider filter that processes multiple measurements over time

	Rover state	Measurements
Observed	\tilde{x}	\tilde{y}
Expected	\hat{x}	\hat{y}

Weighted Batch Filter



$$\text{Minimize } C = \|\tilde{y} - \hat{y}_{k+1}\|_{\mathbf{W}}^2$$

$$\left\{ \begin{array}{l} \mathbf{W} = \text{diag} \left(\sigma_{tot,1}^{-2}, \sigma_{tot,2}^{-2} \right) \\ \sigma_{tot}^2 = \underbrace{\sigma_{e,v}^2}_{\text{Eph. vel. error}} + \underbrace{\sigma_{\dot{\rho}}^2}_{\text{Meas. error}} \end{array} \right.$$

Overview



Measurement Modeling and Filtering

- Simulate and filter realistic measurements from the satellites' communication signals to obtain a positioning fix.



Orbit Design for Auxiliary Satellite

- Determine the orbital elements for the auxiliary satellite that maximizes coverage of the lunar South Pole region.



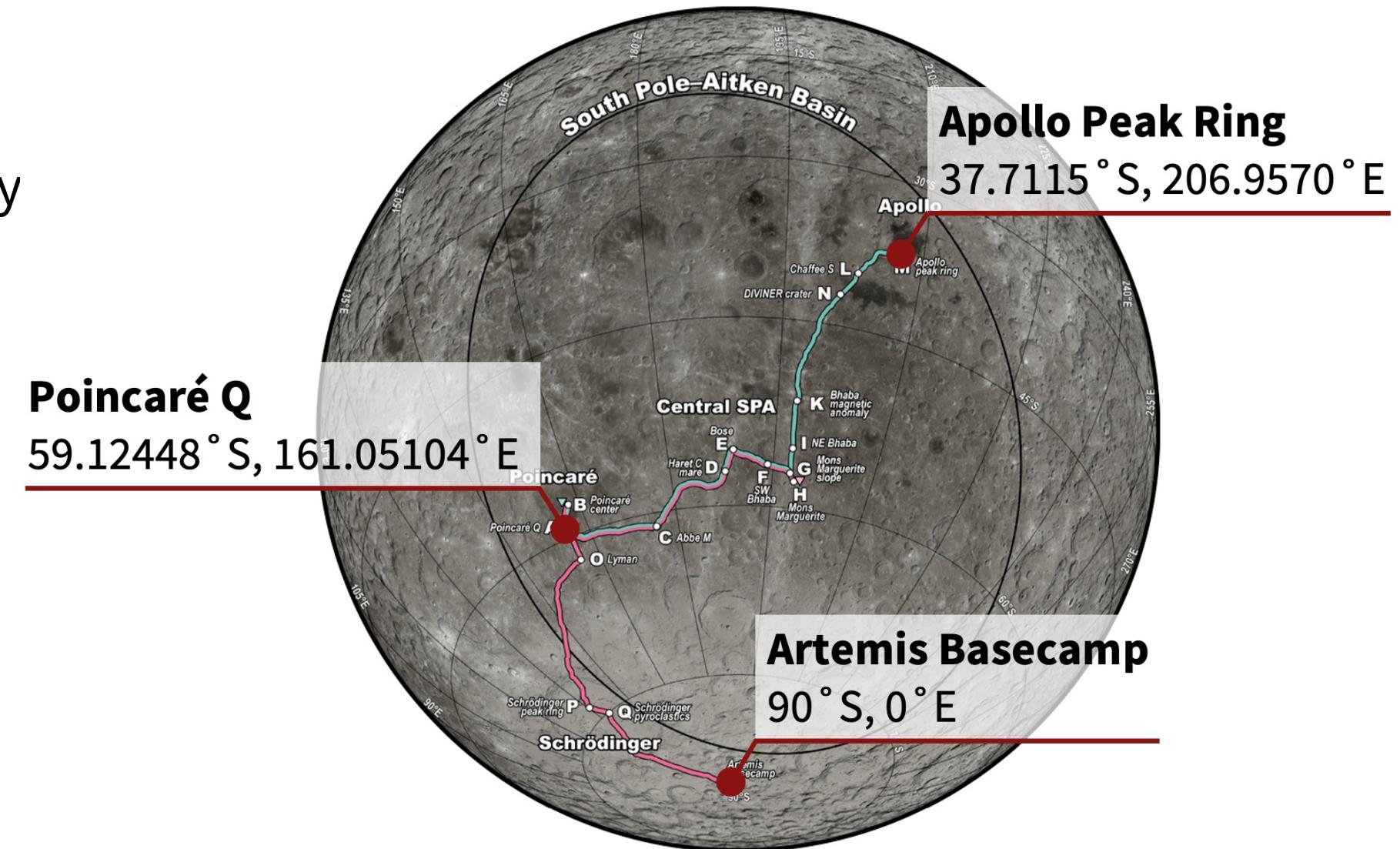
State Estimation Performance and Comparison

- Evaluate the rover's absolute localization accuracy using two non-navigation satellites and compare to the single-satellite scenario.

Rover Positioning Assumptions and Test Sites



- Assume that the rover is stationary during the measurement accumulation window
- Set 3D initial rover positioning error to be 100 m
- Choose three key waypoints as simulation testing locations



Chosen locations along rover's traverse^[5]

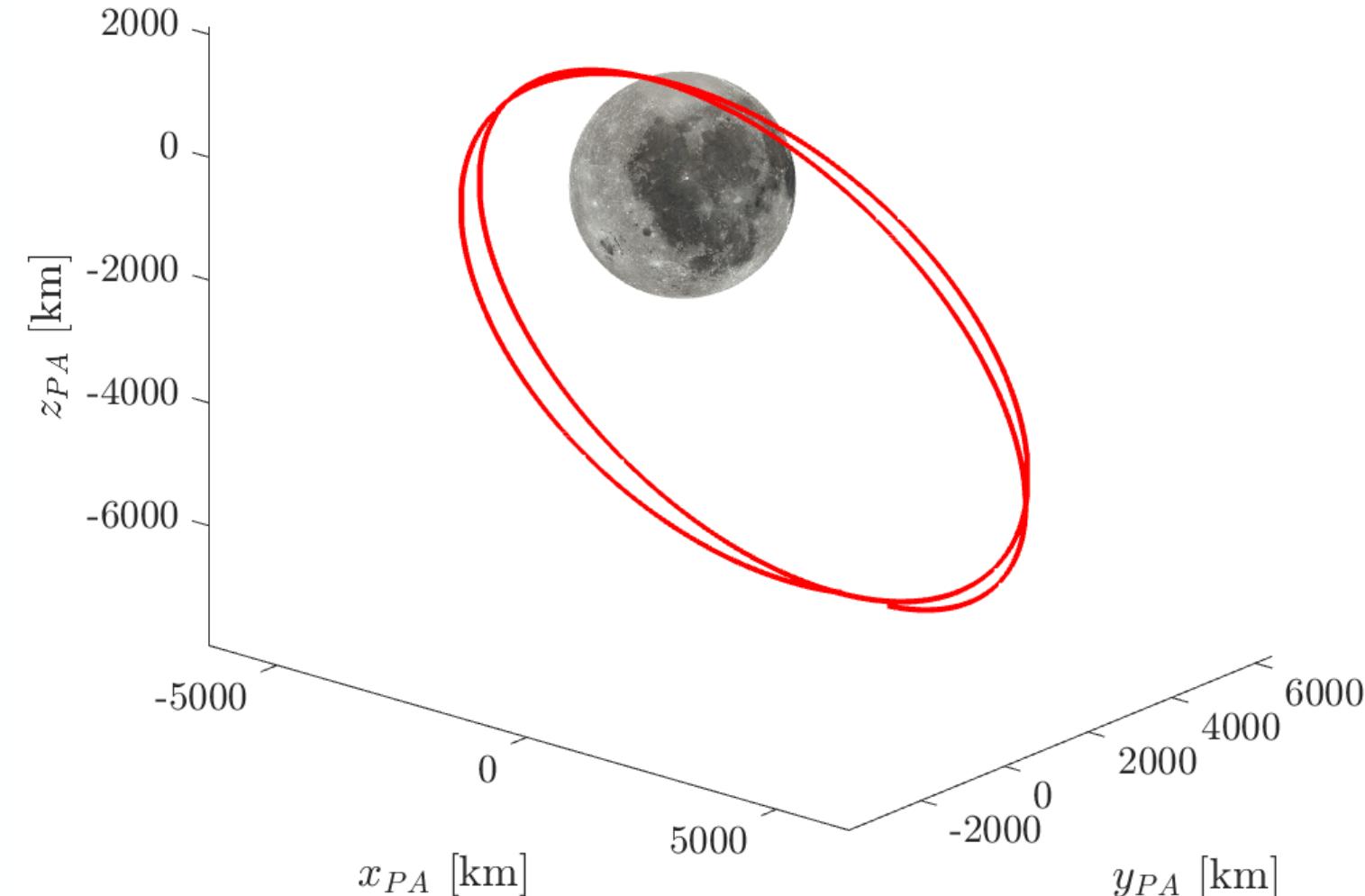
Lunar Pathfinder Orbit Model

Simulation Parameters

LP orbital period	10.84 hours
Initial epoch	2030/10/01 00:00:00 UTC

Lunar Pathfinder's Orbital Elements^[6]

Semi-major axis a	5740 km
Eccentricity e	0.58
Inclination i	54.856 °
RAAN Ω	0 °
Argument of the Periapsis ω	86.322 °
Mean Anomaly M	180 °



Two orbital periods of the Lunar Pathfinder in the fixed Moon Principal Axis frame of reference

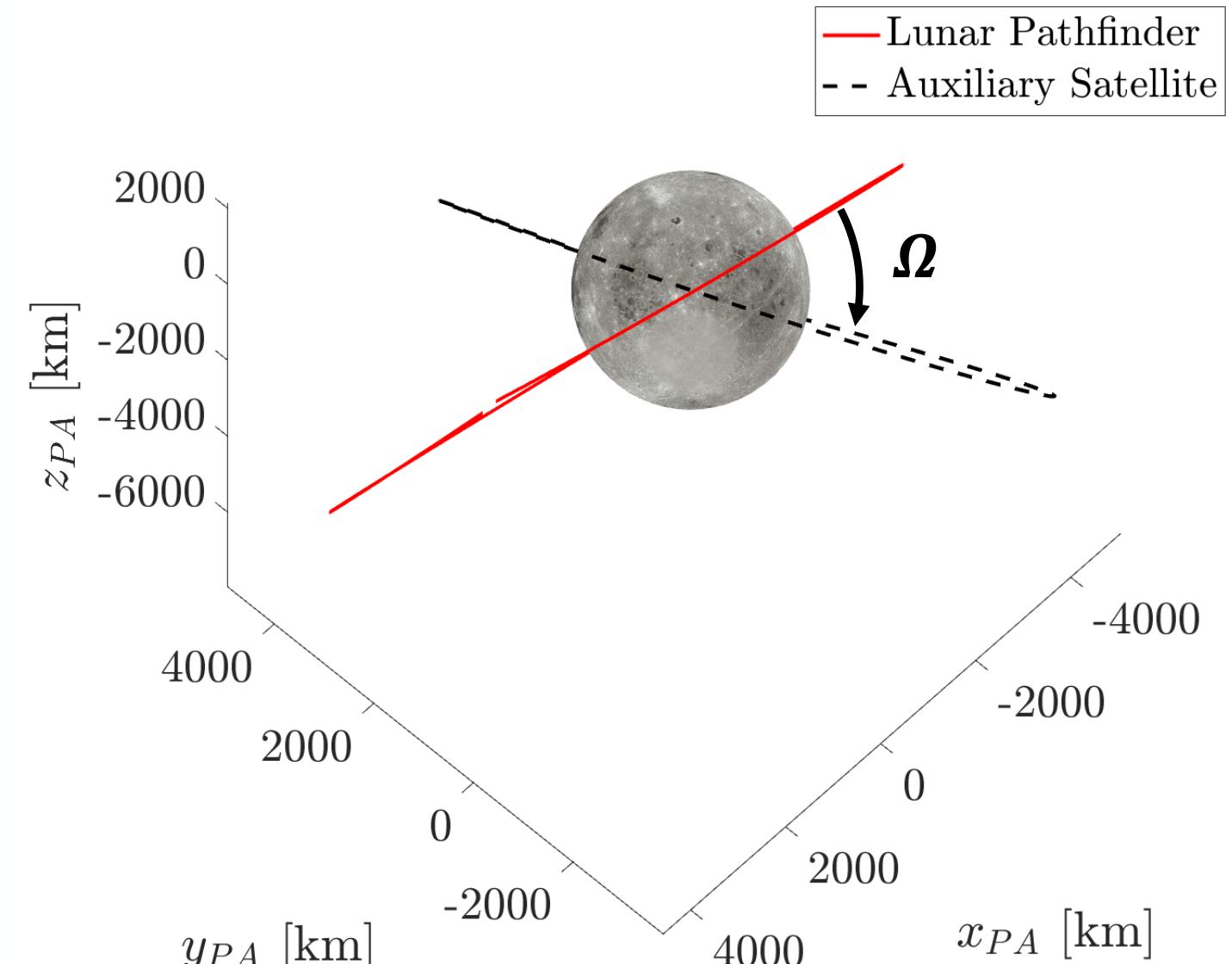
Design Constraints for the Auxiliary Satellite

Goal: Maintain long-duration coverage of the Lunar South Pole region → ELFO

- Freeze a, e, i, ω
- Tune RAAN: Minimize PDOP

Auxiliary Satellite's Orbital Elements

Semi-major axis a	5740 km
Eccentricity e	0.58
Inclination i	54.856°
RAAN Ω	30° to 240°
Argument of the Periapsis ω	86.322°
Mean Anomaly M	180°



Orbits of Lunar Pathfinder and auxiliary satellite in the Moon PA frame (bottom view)

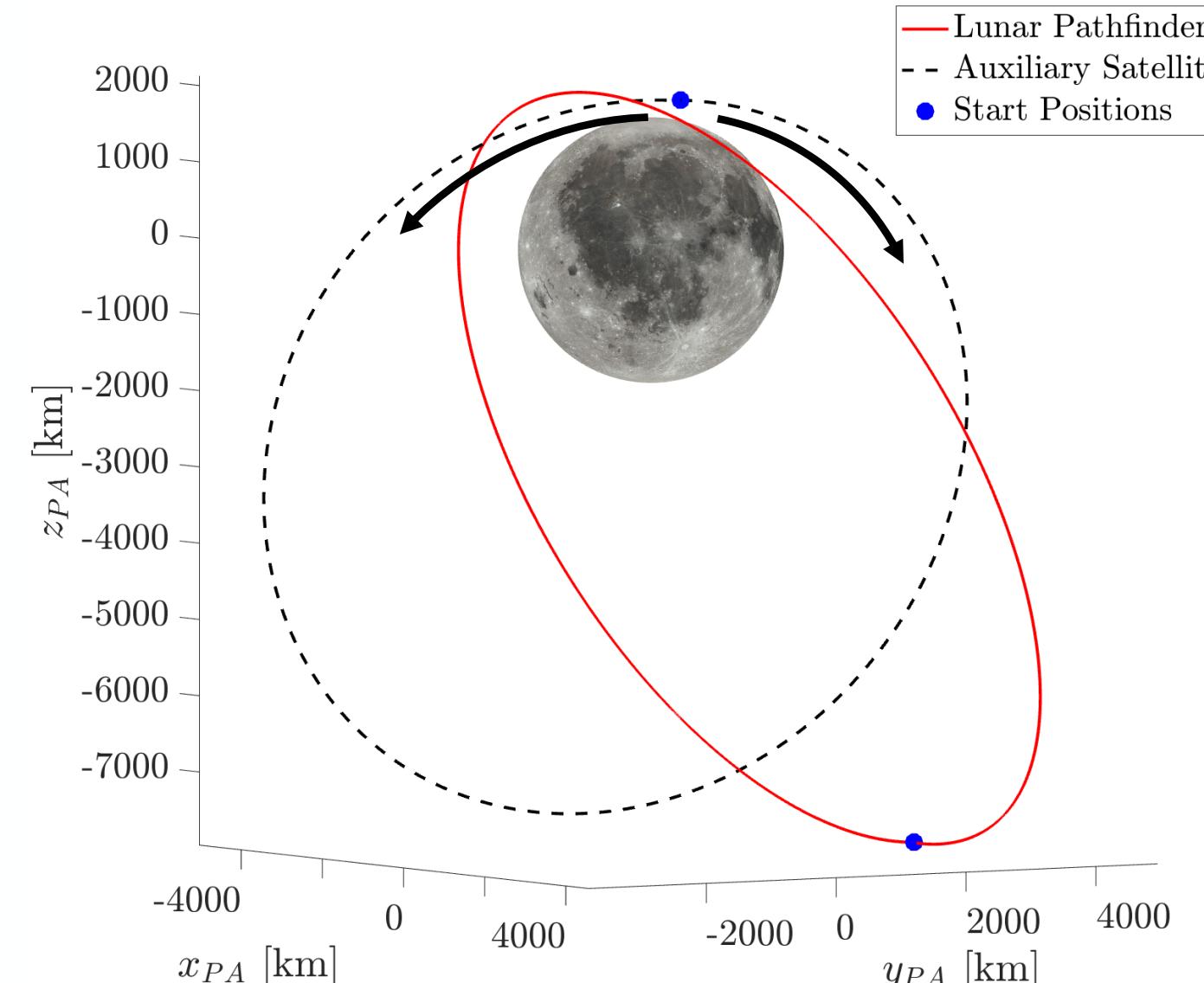
Design Constraints for the Auxiliary Satellite

Goal: Maintain long-duration coverage of the Lunar South Pole region → ELFO

- Freeze a, e, i, ω
- Tune RAAN: Minimize PDOP
- Tune mean anomaly: Maximize visibility

Auxiliary Satellite's Orbital Elements

Semi-major axis a	5740 km
Eccentricity e	0.58
Inclination i	54.856°
RAAN Ω	30° to 240°
Argument of the Periapsis ω	86.322°
Mean Anomaly M	-90° (270°) to 90°

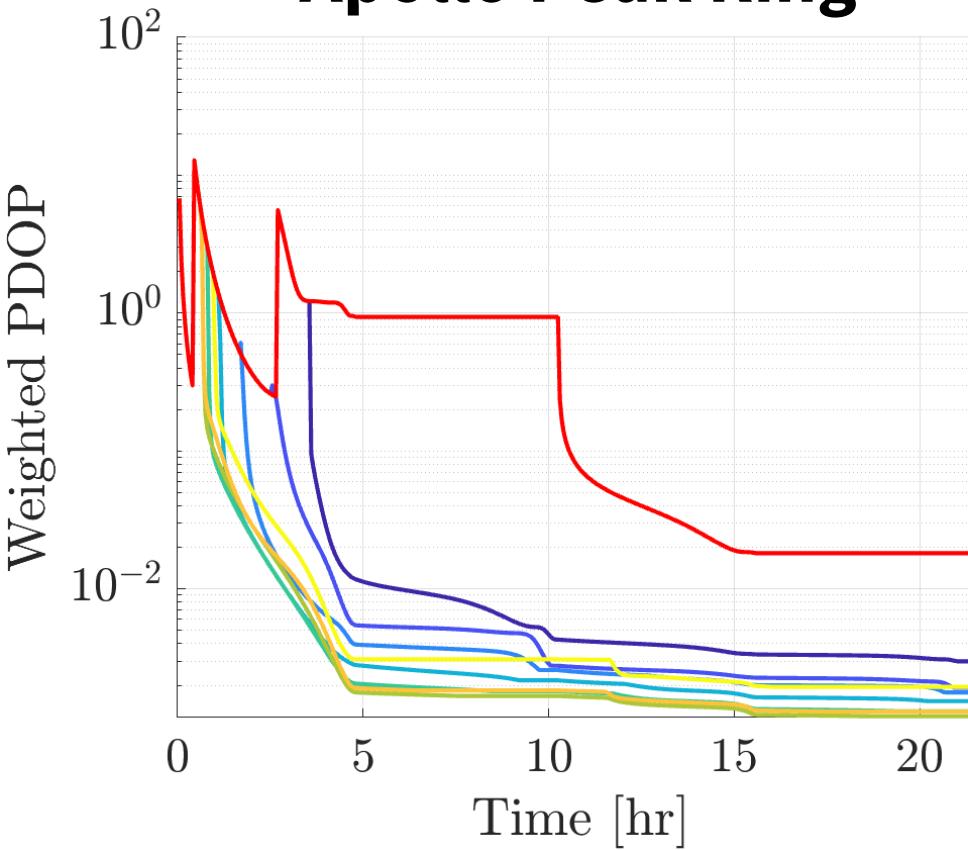


Mean anomaly offset with respect to Lunar Pathfinder's mean anomaly of 180°

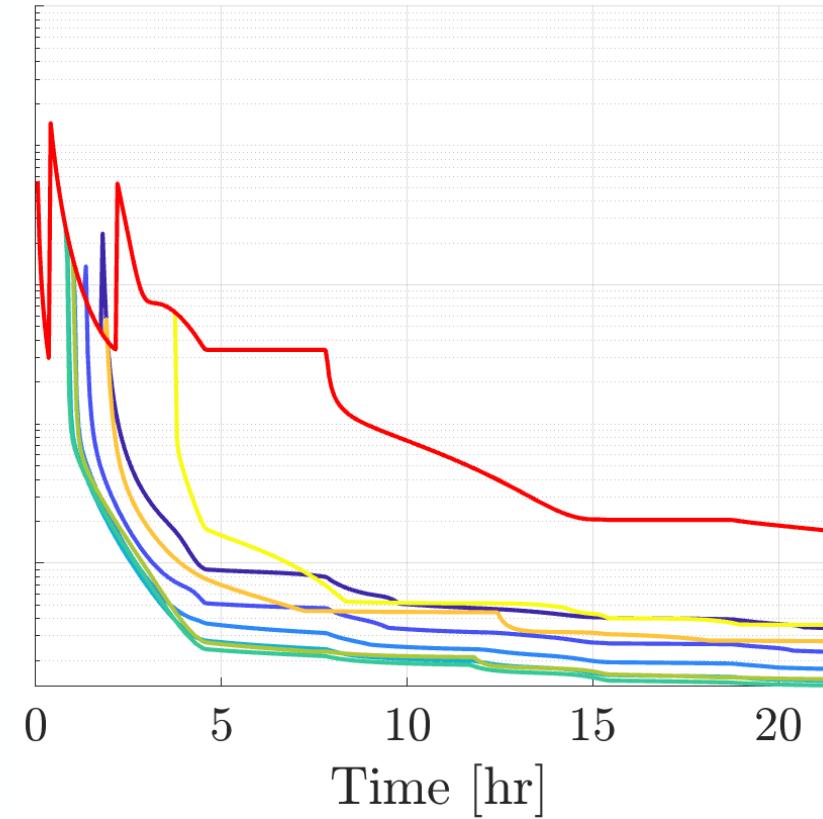
Auxiliary Satellite Orbit Selection: RAAN Ω



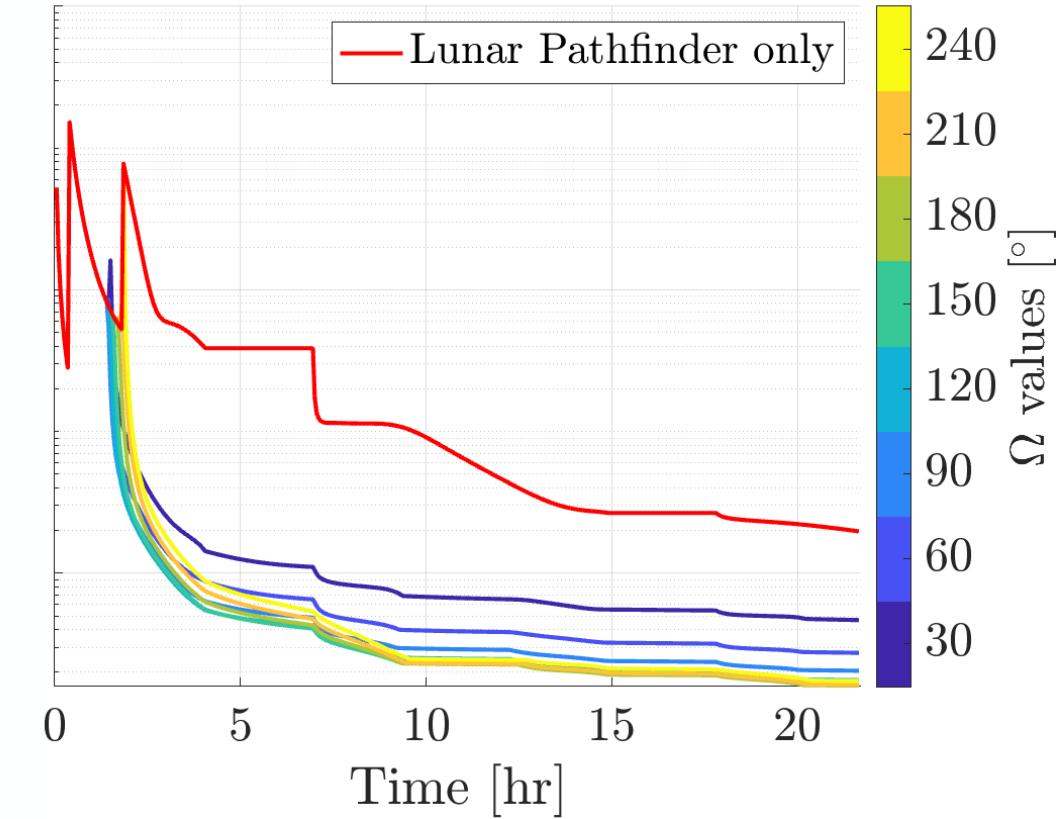
Apollo Peak Ring



Poincaré Q



Artemis Basecamp

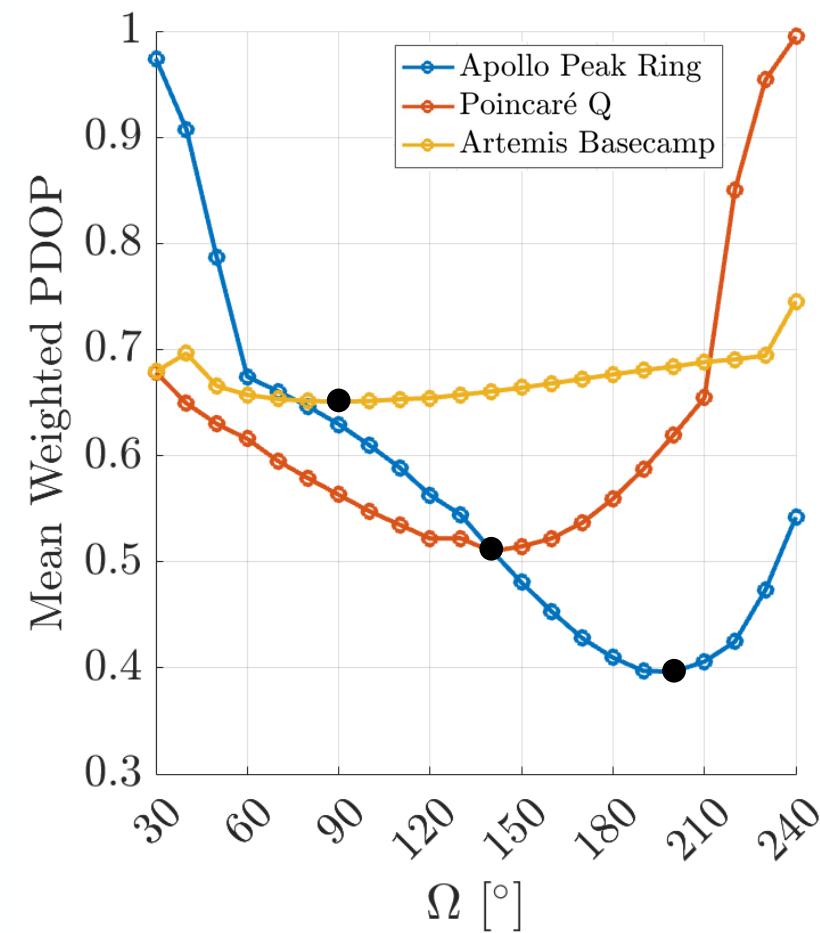


It is not immediately obvious which RAAN value corresponds to the lowest weighted PDOP (WPDOP) across all three locations over time.

Auxiliary Satellite Orbit Selection: RAAN Ω



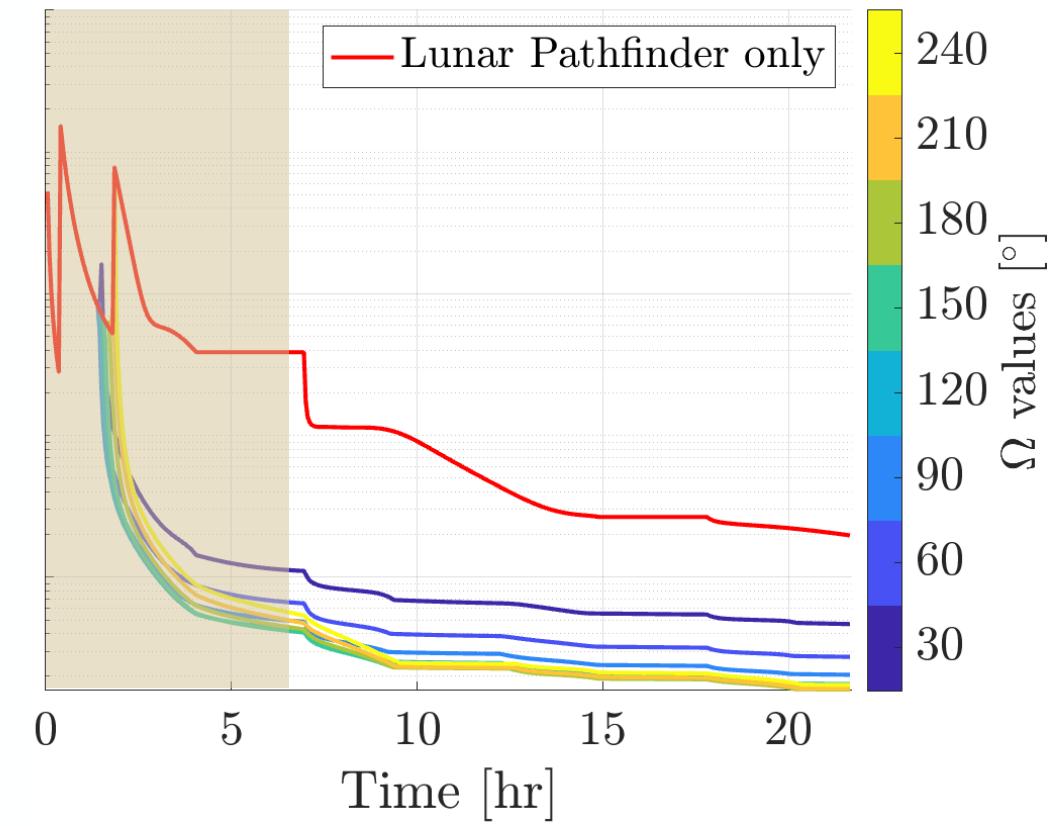
0 - 7 hours



Compute block average over
smaller time segments.

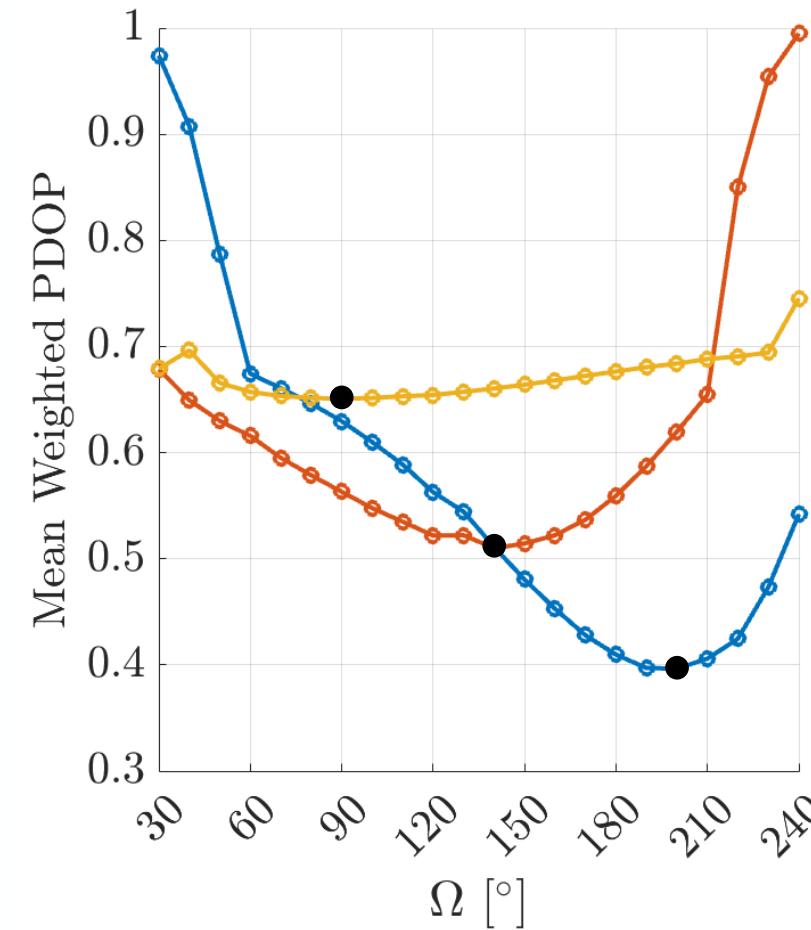


Artemis Basecamp

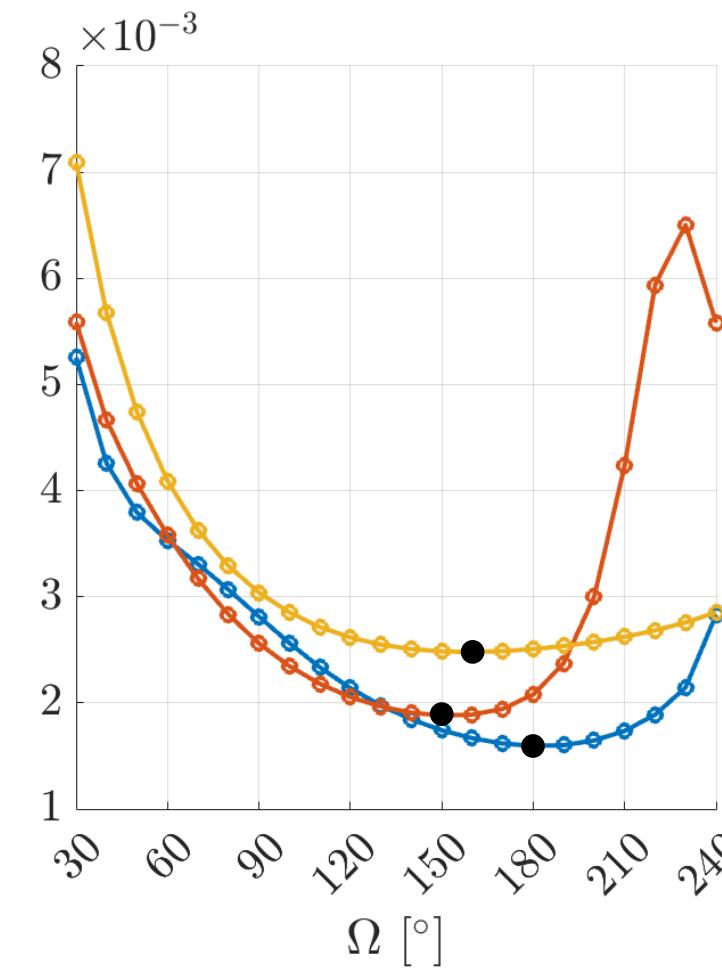


Auxiliary Satellite Orbit Selection: RAAN Ω

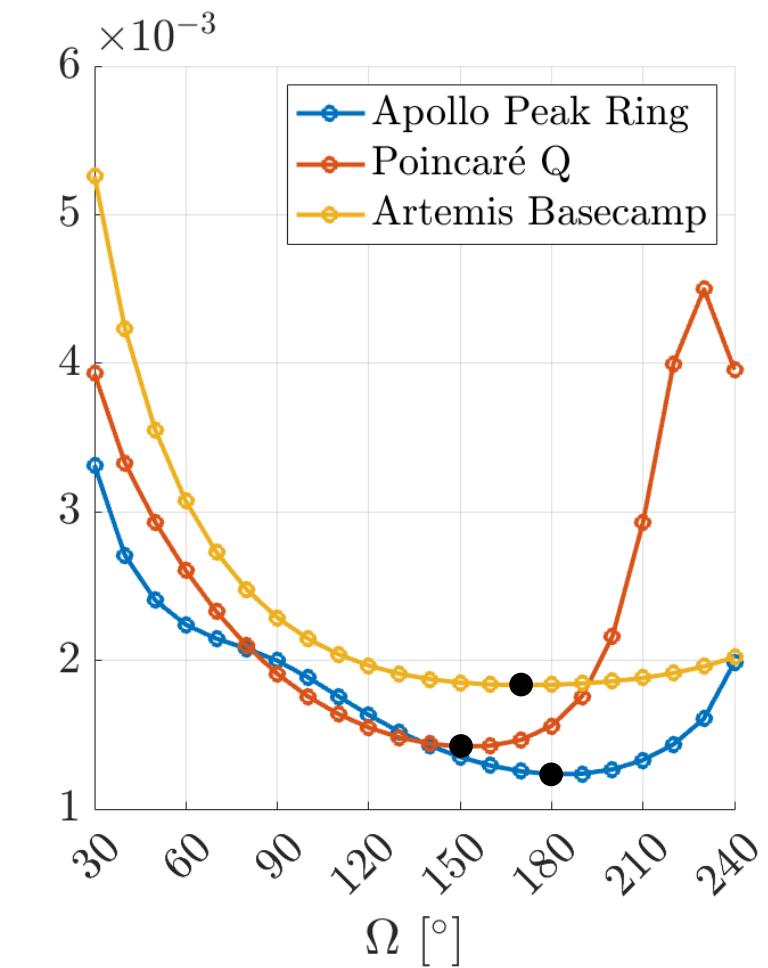
0 - 7 hours



7 - 14 hours



14 - 21 hours



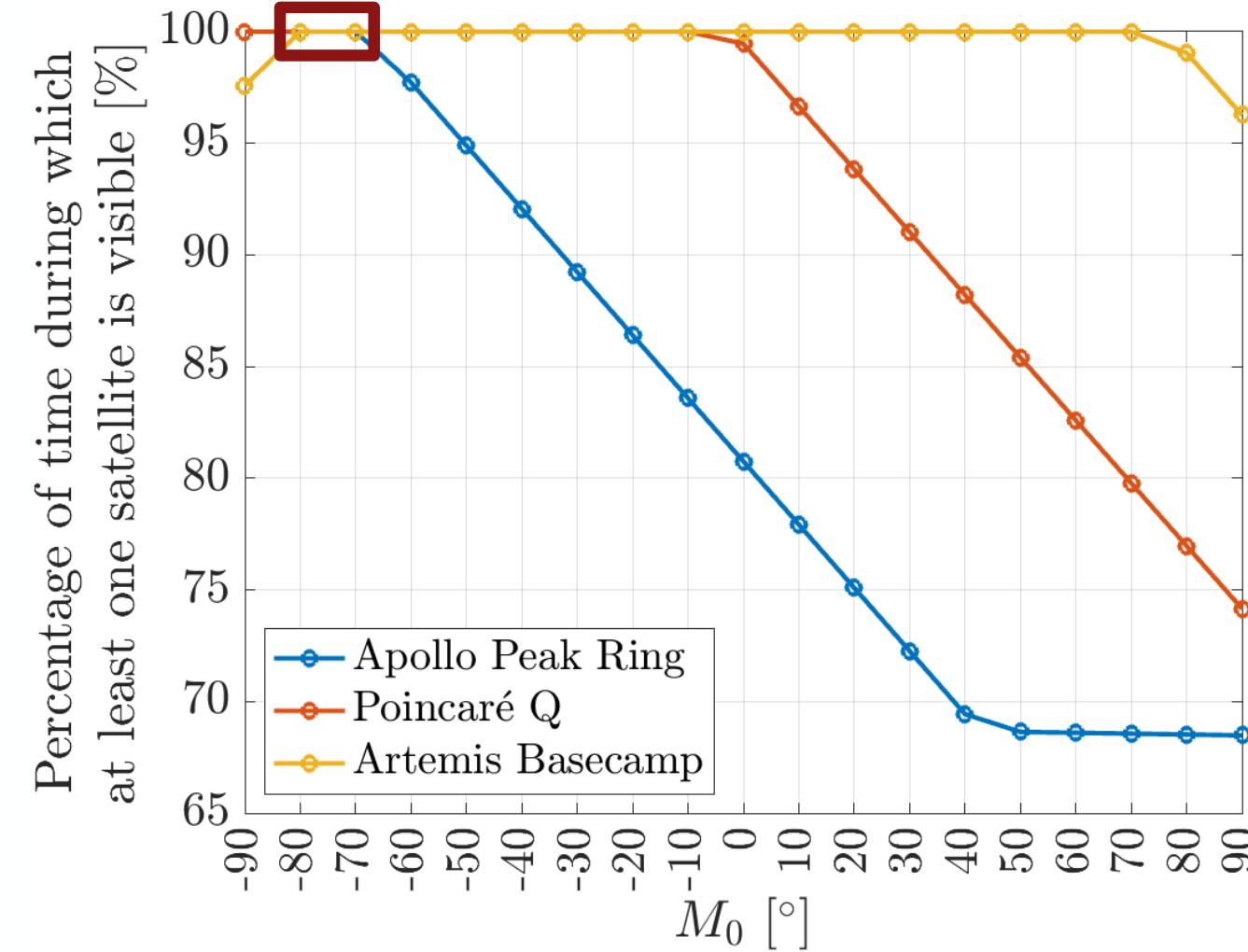
By averaging the lowest WPDOP per time segment across all three locations, the optimal RAAN value is approximately 160° .

Auxiliary Satellite Orbit Selection: Mean Anomaly M_0



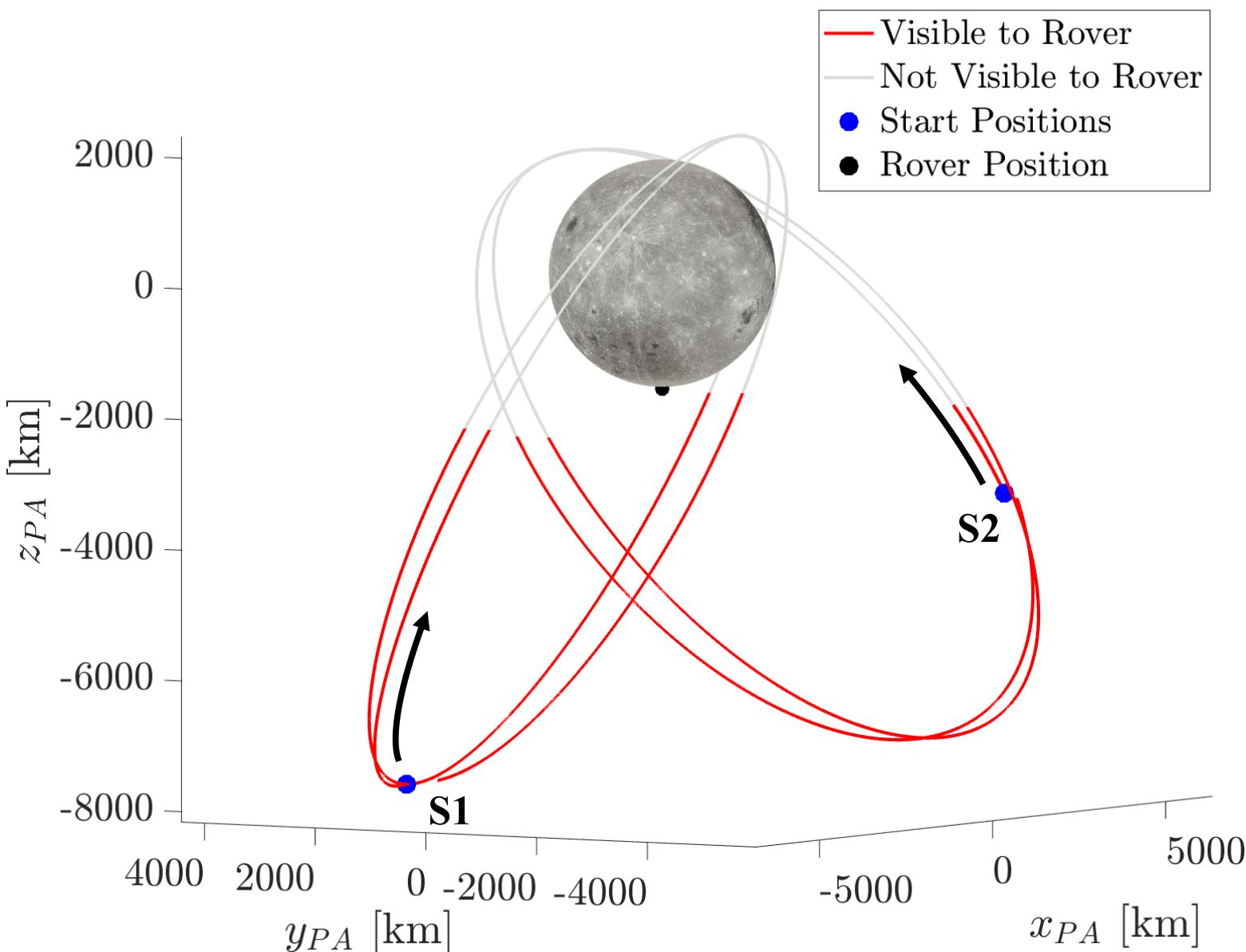
Visibility:

- Elevation mask $> 5^\circ$
- $C/N_0 > 30 \text{ dB Hz}$



To ensure continuous visibility, the mean anomaly of the auxiliary satellite should be -70° (equivalent to 290°), which is a 250° mean anomaly offset relative to the Lunar Pathfinder.

Auxiliary Satellite Orbit Selection



Orbital Elements	Lunar Pathfinder (S1)	Auxiliary Satellite (S2)
Semi-major axis	5740 km	5740 km
Eccentricity	0.58	0.58
Inclination	54.856 °	54.856 °
RAAN	0 °	160 °
Argument of the Periapsis	86.322 °	86.322 °
Mean Anomaly	180 °	290 °

Overview



Measurement Modeling and Filtering

- Simulate and filter realistic measurements from the satellites' communication signals to obtain a positioning fix.



Orbit Design for Auxiliary Satellite

- Determine the orbital elements for the auxiliary satellite that maximizes coverage of the lunar South Pole region.

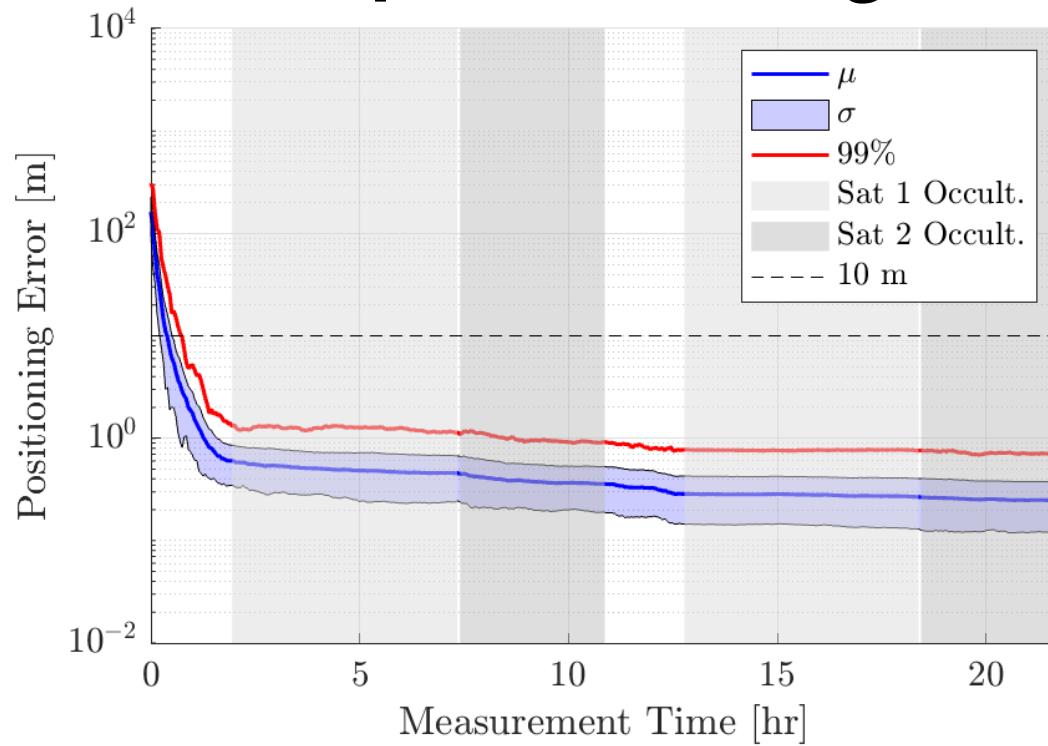


State Estimation Performance and Comparison

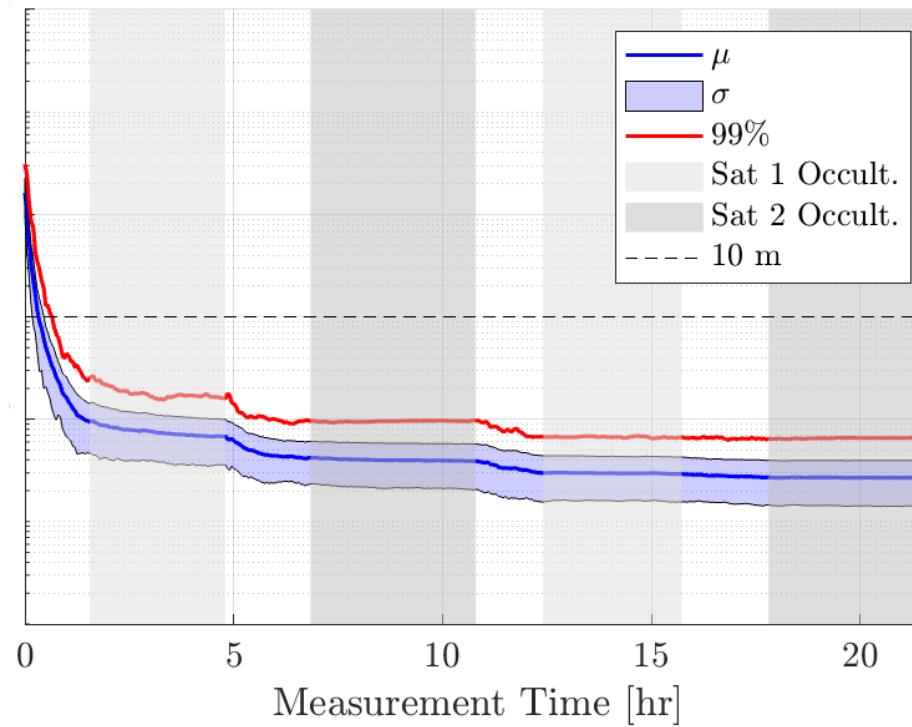
- Evaluate the rover's absolute localization accuracy using two non-navigation satellites and compare to the single-satellite scenario.

Positioning Error across 100 Monte Carlo Runs

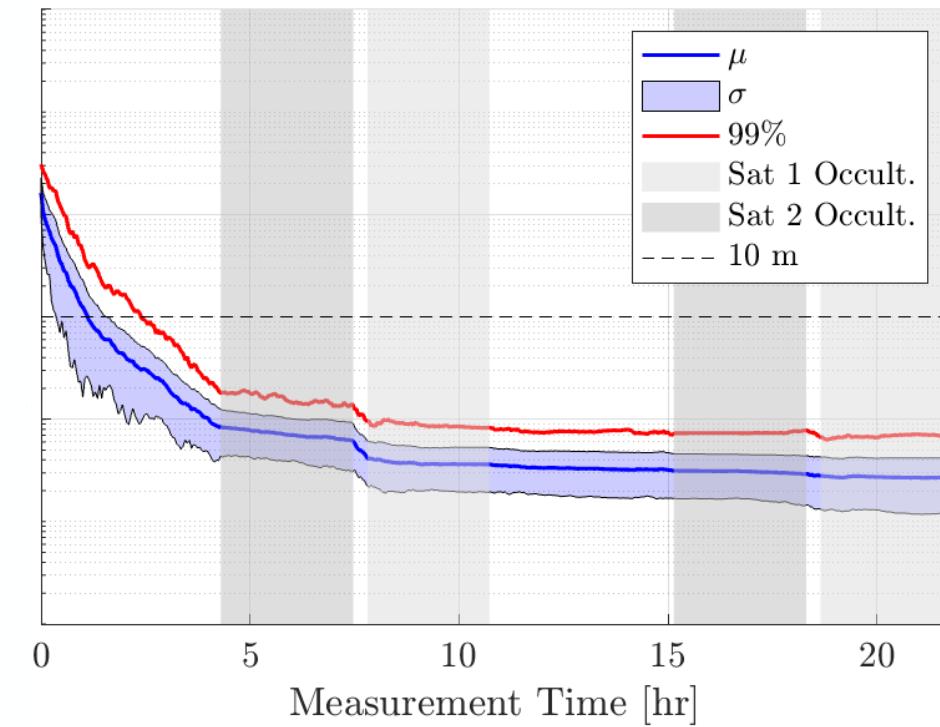
Apollo Peak Ring



Poincaré Q



Artemis Basecamp



Location

Mean [hr]

99% perc. [hr]

Apollo Peak Ring

0.39

0.73

Poincaré Q

0.33

0.65

Artemis Basecamp

1.11

2.39

Key findings:

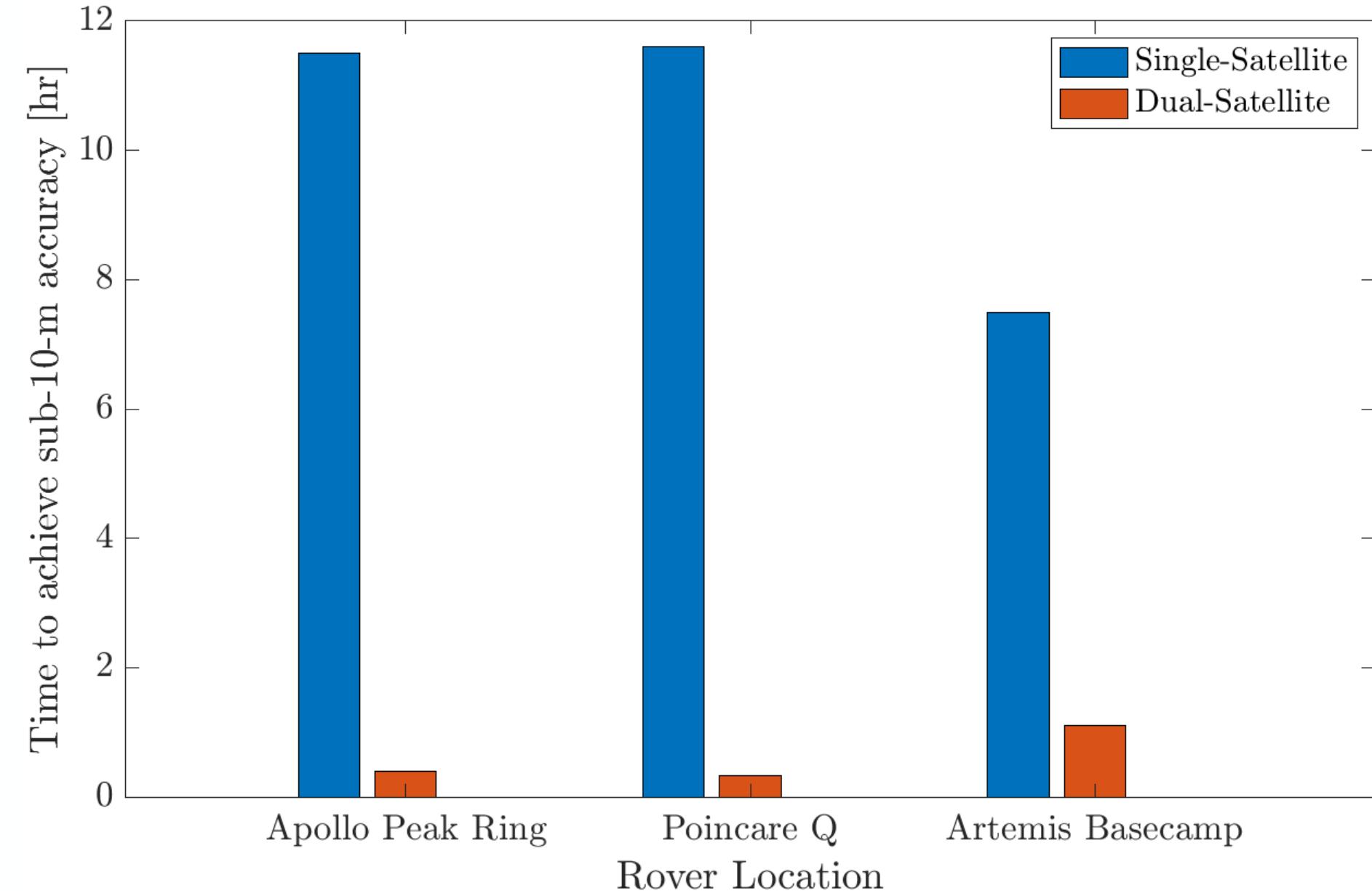
- Mean sub-10-m localization time is ~1 hour
- No overlapping occultation periods

Comparison to the Single-Satellite System

LCRNS ephemeris errors:

- $\sigma_p = 4.48$ m
- $\sigma_v = 0.40$ mm/s

97%, 97%, and 85% improvement with dual-satellite system in comparison to single-satellite system for each location, respectively.



Comparison to the Single-Satellite System

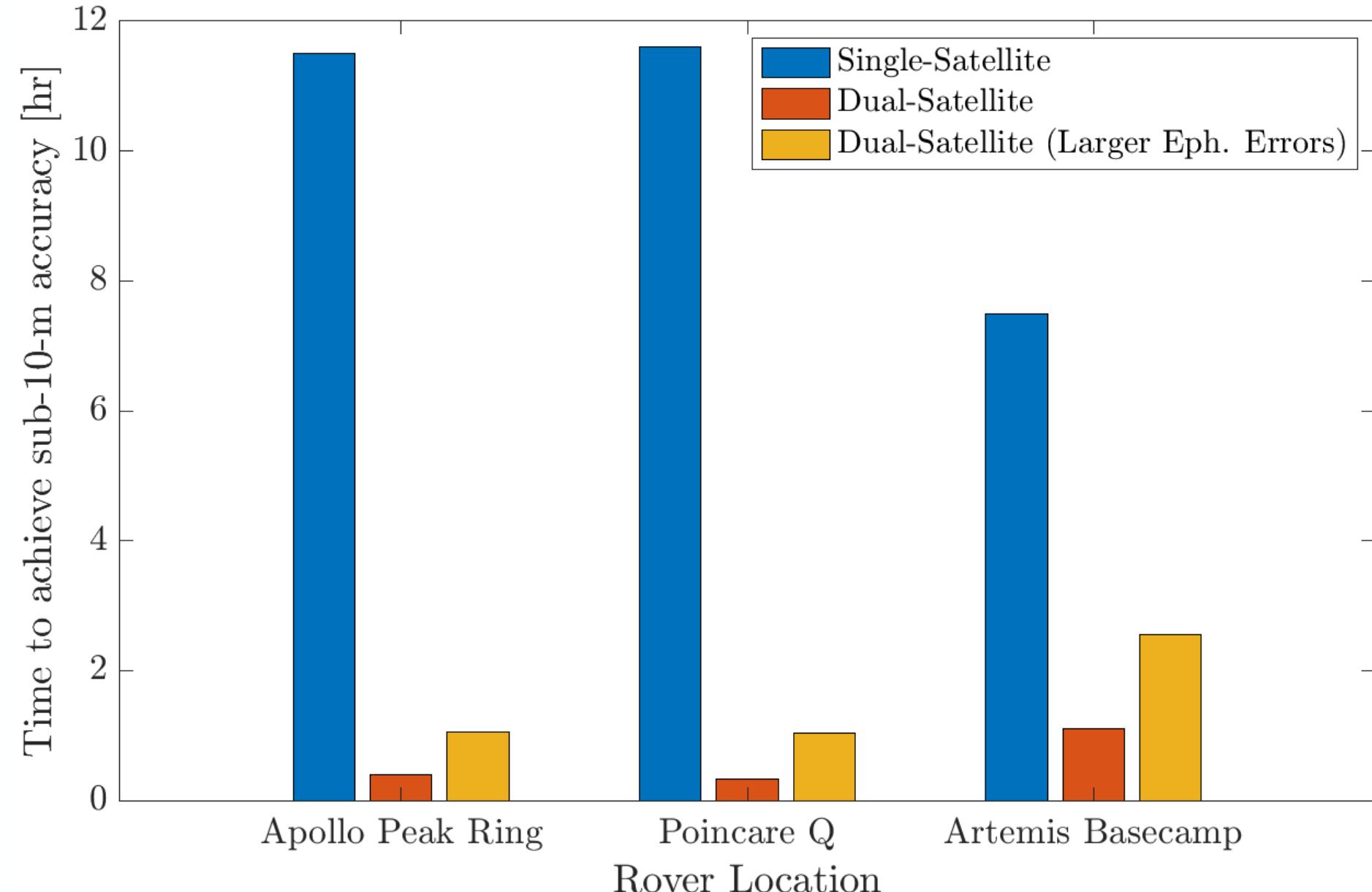
LCRNS ephemeris errors:

- $\sigma_p = 4.48$ m
- $\sigma_v = 0.40$ mm/s

Inflated ephemeris errors:

- $\sigma_p = 100.0$ m
- $\sigma_v = 10.0$ mm/s

91%, 91%, and 66% improvement with inflated eph. errors in comparison to single-satellite system for each location, respectively.



Concluding Remarks

Key Questions

What will the orbit of the auxiliary satellite have to look like to best complement the Lunar Pathfinder's coverage?

What *improvement* in the rover's localization accuracy can be achieved by introducing a second satellite?

Main Results

We identified the optimal auxiliary satellite orbit that minimizes PDOP, maximizes visibility, and maintains an ELFO.

We achieved an 85% to 97% improvement over the single-satellite system when using similar ephemeris errors.

Future Work

- Perform 2D grid search on auxiliary satellite's orbital elements
- Fuse with measurements from the onboard sensors to localize the rover while it is in motion

Acknowledgements



Special thanks to the NAV Lab!

Comparative Analysis and Design of a Dual-Satellite System for Lunar Rover Localization



KAILA M. Y. COIMBRA AND GRACE GAO

A high-resolution image of the Moon in the foreground, showing its craters and surface texture. In the background, the Earth is visible as a blue and green sphere against a dark, star-filled space.

Thank you!

Selected References

1. The Planetary Society, “Every Mission to the Moon, Ever,” *The Planetary Society*, n.d. [Online]. Available: <https://www.planetary.org/space-missions/every-moon-mission>
2. ISRO. “Chandrayaan-3 Details,” *ISRO*, 2023. [Online]. Available: https://www.isro.gov.in/Chandrayaan3_Details.html
3. NASA. “Intuitive Machines 1 (Odysseus),” *NASA*, 2024. [Online]. Available: <https://nssdc.gsfc.nasa.gov/nmc/spacecraft/display.action?id=IM-1-NOVA>
4. The Planetary Society, “Chang'e-6, collecting the first lunar farside samples,” *The Planetary Society*, n.d. [Online]. Available: <https://www.planetary.org/space-missions/change-6>
5. J. T. Keane, S. M. Tikoo, and J. Elliot, “Endurance: Lunar South Pole-Aitken Basin Traverse and Sample Return Rover,” *NASA Jet Propulsion Laboratory, Technical Report*, 2022. [Online]. Available: <https://tinyurl.com/2p88fx4f>
6. SSTL, “Lunar Pathfinder: Data relay satellite in orbit around the Moon,” *SSTL, Service Guide (V4)*, 2022. [Online]. Available: <https://www.sstl.co.uk/getmedia/9efe3cc5-6b5b-4fd1-bbd7-443bc2d21848/Lunar-Pathfinder-Service-Guide-V004.pdf>
7. M. Cortinovis, T. Mina, and G. Gao, “Assessment of Single Satellite-based Lunar Positioning for the NASA Endurance Mission,” *2024 IEEE Aerospace Conference*, 2024.
8. K. M. Y. Coimbra, M. Cortinovis, T. Mina, and G. Gao, “Single-Satellite Lunar Navigation via Doppler Shift Observables for the NASA Endurance Mission,” *Proceedings of the Institute of Navigation GNSS+ conference (ION GNSS+ 2024)*, 2024.
9. B. L. Schmittberger and D. R. Scherer, “A Review of Contemporary Atomic Frequency Standards,” *arXiv preprint arXiv:2004.09987*, 2020.
10. S. Bhamidipati, T. Mina, and G. Gao, “A Case Study Analysis for Designing a Lunar Navigation Satellite System with Time-Transfer from Earth-GPS,” *Proceedings of the Institute of Navigation ITM conference (ION ITM 2022)*, 2022.

State Estimation Framework

Weighted Batch Filter

Accumulate measurements.

$$\tilde{\mathbf{y}} = [\tilde{\rho}_1(t_1) \ \dots \ \tilde{\rho}_1(t_N) \ \tilde{\rho}_2(t_1) \ \dots \ \tilde{\rho}_2(t_N)]^\top$$

Predict the next expected pseudorange rate measurements.

$$\begin{aligned}\hat{\dot{\rho}}_{k+1}(t) &= \tilde{\mathbf{v}}_s(t) \cdot \frac{\tilde{\mathbf{r}}_s(t) - \hat{\mathbf{r}}_{r,k+1}}{\|\tilde{\mathbf{r}}_s(t) - \hat{\mathbf{r}}_{r,k+1}\|} + c \left(\hat{\delta t}_{r,k+1} - \hat{\delta t}_s \right) \\ \hat{\mathbf{y}}_{k+1} &= [\hat{\dot{\rho}}_{1,k+1}(t_1) \ \dots \ \hat{\dot{\rho}}_{2,k+1}(t_N)]^\top\end{aligned}$$

Minimize cost function.

$$\begin{aligned}C &= \|\tilde{\mathbf{y}} - \hat{\mathbf{y}}_{k+1}\|_{\mathbf{W}}^2 \\ \delta \mathbf{x}_k &= (\mathbf{J}_k^\top \mathbf{W} \mathbf{J}_k)^{-1} \mathbf{J}_k^\top \mathbf{W} \delta \mathbf{y}_k \quad , \quad \mathbf{W} = \text{diag} \left(\sigma_{tot,1}^{-2}, \sigma_{tot,2}^{-2} \right)\end{aligned}$$

Simulation Parameters

Measurement sampling rate	1 Hz
Filter update interval	180 seconds
Number of Monte Carlo runs	100

Use weighted PDOP to determine measurement diversity.

$$\begin{aligned}\mathbf{M} &= (\mathbf{J}_k^\top \mathbf{W} \mathbf{J}_k)^{-1} \\ PDOP_{\mathbf{W}} &= \sqrt{M_{11} + M_{22} + M_{33}}\end{aligned}$$

State Estimation Framework

- We can remove the initial spike by augmenting the measurement vector with the rover's initial position estimate
- Doing so does not compromise the long-term performance of the state estimation

